

# The FORMATION of TOPOLOGICAL DEFECTS

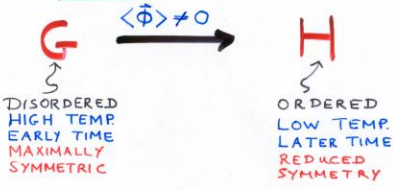
- M.B. L. CHANDAR, E. SCHIFF & A. SRIVASTAVA :  
SCIENCE 263 (1994) 943 [hep-ph/9208293](#)
- M.B. & A. MOMEN,  
PHYS. REV. D58 (1998) 085014 [hep-ph/9803284](#)
- M.B. & A. CACCIUTO, A. TRAVESSET  
[Cond-mat/0107188](#)

## KEY INGREDIENTS

- SPONTANEOUS SYMMETRY BREAKING (SSB)
- TOPOLOGICALLY NON-TRIVIAL VACUUM / GROUND STATE MANIFOLD  $\mathcal{M}$
- CONTINUOUS PHASE TRANSITION OR FIRST ORDER TRANSITION PROCEEDING VIA SPINODAL DECOMPOSITION

- FINITE QUENCH  
RATE / EXPANSION  
RATE OF UNIVERSE

▲ SSB



▲ VACUUM MANIFOLD

$$\mathcal{M} = G/H$$

NON-TRIVIAL TOPOLOGY for  $\mathcal{M}$   
 $\Rightarrow$  TOPOLOGICAL DEFECTS

Exs: 1.  $\pi_0(\mathcal{M}) \neq 0 \Leftrightarrow$  DOMAIN WALLS

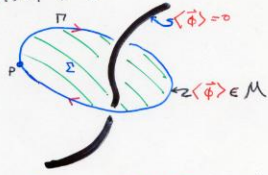
i.e.  $\mathcal{M}$  IS NOT CONNECTED

e.g.  $\mathcal{M} = \mathbb{Z}_2$   
 (ISING)



dim(DEFECT) = 2

2.  $M$  IS NOT SIMPLY-CONNECTED  
 $\pi_1(M) \neq 0$



$\Leftrightarrow$  LINE/LOOP/STRING DEFECT ( $d=1$ )

EXAMPLES:

- COSMIC STRINGS
- SUPERFLUID  $He^4$  VORTEX LINES
- SUPERCONDUCTING FLUX TUBES
- NEMATIC LIQUID CRYSTAL DISCLINATIONS
- SUPERFLUID  $He^3$  VORTEX LINES

3.  $\pi_2(\mathcal{M}) \neq 0$

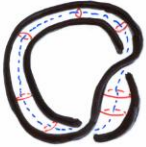
$\Leftrightarrow$  POINT (MONOPOLE)  
DEFECT



dim = 0

4.  $\pi_3(\mathcal{M}) \neq 0$

$\Leftrightarrow$  TEXTURE



dim = 3  
"SKYRMION"

## GLOBAL COSMIC STRINGS

•  $G = U(1) \rightarrow 1$

$$M \cong S^1 \quad \pi_1(S^1) \cong \mathbb{Z}$$

•  $G = SO(10) \xrightarrow{126} SU(5) \times \mathbb{Z}_2$

$$M \cong SO(10)/SU(5) \times \mathbb{Z}_2$$

$$\pi_1(M) = \mathbb{Z}_2$$

FORM AT  $t \sim 10^{-37} \text{ s}$  AFTER

BIG BANG  $T \sim 10^{28} \text{ K}$

$\mu \sim 10^{19} \text{ kg/cm}$

PLANAR SPINS

$\mathcal{M} = S^1 \quad \pi_1(S^1) = \mathbb{Z} \Rightarrow$  VORTICES



ORDINARY SPINS

$\mathcal{M} = S^2 \quad \pi_2(S^2) = \mathbb{Z} \Rightarrow$  MONOPOLES

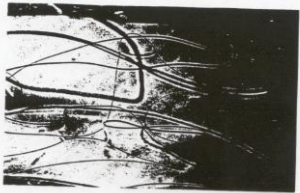


NEMATICS

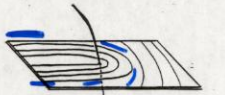
$\mathcal{M} = S^2/\mathbb{Z}_2 \cong \mathbb{R}P^2 \cong O(3)/D_{\text{oh}}$

$\pi_1(\mathcal{M}) = \mathbb{Z}_2$     $\pi_2(\mathcal{M}) = \mathbb{Z}/\mathbb{Z}_2$     $\pi_3(\mathcal{M}) = \mathbb{Z}$   
DISCLINATIONS   MONOPOLES   TEXTURE



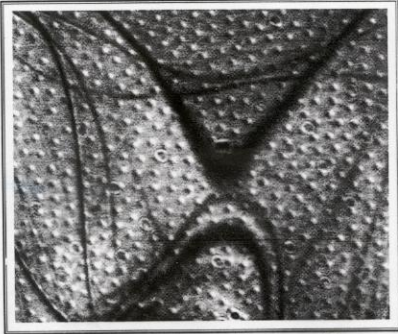


LINE DEFECTS IN  
NEMATICS (M. KLÉMAN)



# BULLETIN

OF THE AMERICAN PHYSICAL SOCIETY



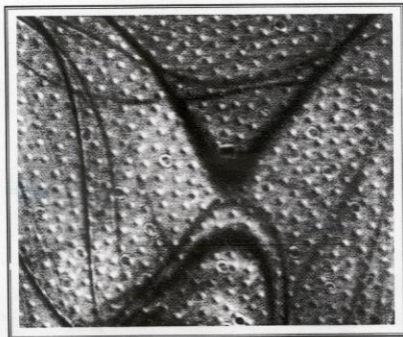
LORENTZ MICROGRAPH OF Nb FILM ( $B=100G$ )  
(HITACHI ADVANCED RESEARCH LAB, JAPAN)

Program of the 1994 March Meeting

March 1994  
Volume 39, No. 1

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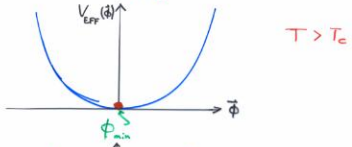
# CONTINUOUS PHASE TRANSITION

COMPUTE  $V_{\text{EFF}}(\vec{\phi})$

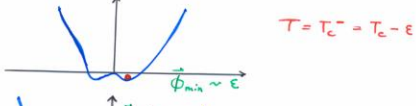
FOR  $T > T_c$   $\langle \vec{\phi}_{\text{min}} \rangle = 0$  DISORDERED

$T < T_c$   $\langle \vec{\phi} \rangle \neq 0$  ORDERED

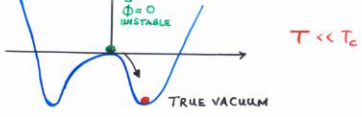
$\langle \vec{\phi} \rangle_{T=T_c} \equiv 0$  CRITICAL POINT



$T > T_c$

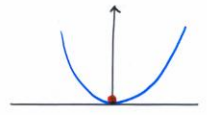


$T = T_c - \epsilon$

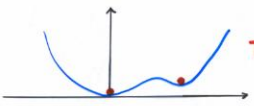


$T \ll T_c$

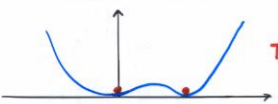
FIRST OR



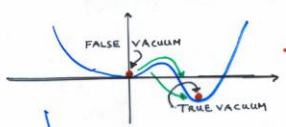
$T \gg T^*$



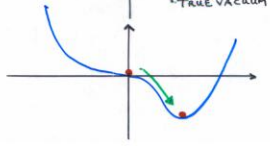
$T = T^*$



$T = T^*$



$T = (T^*)'$

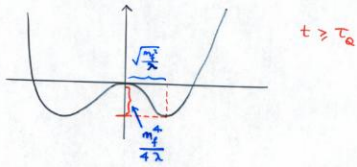


$T = T_{SPINDAL}$

Let  $\epsilon = \frac{T - T_c}{T_c}$  REDUCED TEMPERATURE

$m^2 = m_0^2 \epsilon$  IN VICINITY OF TRANSITION

$\lambda = \text{CONST.}$



$$|\dot{\phi}_{\min}| = v^2 = m_0^2 / \lambda$$

$$\Delta V = V(0) - V(\phi_{\min}) = m_0^4 / 4\lambda$$

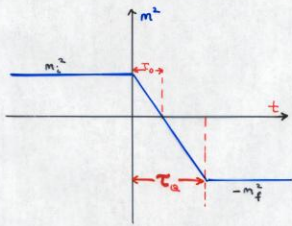
# MODEL PHASE TRANSITION VIA

## LANDAU-GINZBURG EFFECTIVE POTENTIAL / FREE ENERGY

$$V_{eff}(\vec{\phi}) = \frac{1}{2} |\nabla \vec{\phi}|^2 - \frac{1}{2} m^2(t) |\vec{\phi}|^2 - \frac{\lambda}{4} |\vec{\phi}|^4$$

with

$$m^2(t) = \begin{cases} m_i^2 & t \leq 0 \\ m_i^2 - \left(\frac{m_i^2 + m_f^2}{\tau_a}\right) t & 0 \leq t \leq \tau_a \\ -m_f^2 & t \geq \tau_a \end{cases}$$



$$m_i^2 = \mu^2 \left( \frac{T_i}{T_c} - 1 \right)$$

$$m_f^2 = \mu^2 \left( 1 - \frac{T_f}{T_c} \right)$$



E.O.M :

$$\left[ \frac{d^2}{dt^2} + \vec{k}^2 + m_i^2 \right] U_k(t) = 0 \quad t \leq 0$$

SINGLE-PARTICLE  
WAVE FUNCTION

$$U_k(t) = e^{-i\omega(k)t} \quad \omega^2(k) = \vec{k}^2 + m_i^2$$

INITIAL CONDITION

$T_q \rightarrow 0$  SUDDEN QUENCH

$$\left[ \frac{d^2}{dt^2} + \vec{k}^2 - m_f^2 \right] U_k(t) = 0 \quad t \geq 0$$

LONG-WAVELENGTH MODES w/  $\vec{k}^2 < m_f^2$

ARE UNSTABLE (SPINODAL)

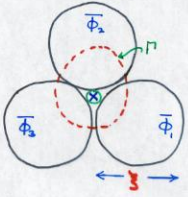
⇒ FORMATION OF ORDERED DOMAINS

GUTH & PI (1985) E. WEINBERG & A. WU (1989)

$$U_k(t) = A_k e^{W(k)t} + B_k e^{-W(k)t}$$

$$w/ \quad W(k) = \sqrt{m_f^2 - \vec{k}^2}$$

DOMAIN FORMATION  
A LA KIBBLE (1976, 1980)



ASSUMPTIONS

- $\bar{\phi}_i$  UNIFORM WITHIN DOMAIN OF SIZE  $\xi$
- $\bar{\phi}_i$  RANDOMLY CHOSEN FROM  $\mathcal{M}$   
i.e. DOMAINS ARE UNCORRELATED
- ALONG PATH  $\Gamma$  ORDER PARAMETER  $\langle \phi \rangle$  SMOOTHLY INTERPOLATES FROM

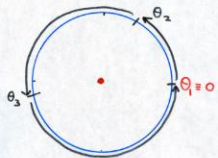


= **GEODESIC RULE**

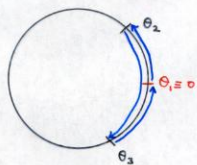
HORIZON SIZE (COSMOLOGY)

$\xi_{\text{max}}$  MEAN CORRELATION LENGTH (STAT. MECH.)

EX.  $M = S^1 \times \mathbb{R}^2$



+1 VORTEX  
NON-CONTRACTIBLE  
LOOP FROM  
GEODESIC RULE

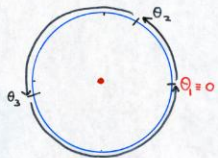


NO VORTEX  
CONTRACTIBLE  
LOOP

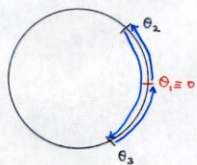
VORTEX IFF  $\theta_3 > \pi$  &  $|\theta_3 - \theta_2| < \pi$

$$\text{PROBABILITY} = \frac{1}{2} \times \frac{1}{2} = \frac{1}{4}$$

EX.  $M = S^1 \times \mathbb{R}^2$



+1 VORTEX  
NON-CONTRACTIBLE  
LOOP FROM  
GEODESIC RULE



NO VORTEX  
CONTRACTIBLE  
LOOP

VORTEX IFF  $\theta_3 > \pi$  &  $|\theta_3 - \theta_2| < \pi$

$$\text{PROBABILITY} = \frac{1}{2} \times \frac{1}{2} = \frac{1}{4}$$

$$n_{\text{defects}} = \frac{\# \text{ DEFECTS}}{\text{VOLUME}} \\ \sim \frac{1}{\xi^*} d \quad d = \text{dimension}$$

$$\xi^* = \xi \text{ (FORMATION)}$$

Q: WHAT DETERMINES  $\xi^*$

A: KIBBLE (1980), ZUREK (93)

ORDER FIELD(S) CAN EQUILIBRATE  
(REACH A UNIFORM VACUUM) PROVIDED  
THE INTRINSIC DYNAMICS IS FASTER  
THAN THE QUENCH RATE

$$\Gamma_{\text{in}} \geq \Gamma_{\text{q}}$$

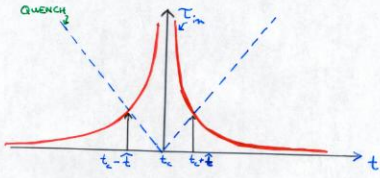
For  $\Gamma_{\text{in}} < \Gamma_{\text{q}}$  ORDERING IS  
FROZEN

But (!) IN A CONTINUOUS TRANSITION

THE INTRINSIC DYNAMICS BECOMES  
ARBITRARILY **SLUGGISH**  
AS  $T \rightarrow T_c$

Take  $\epsilon = \frac{T - T_c}{T_c} = t / \tau_a$  **QUENCH**

$\tau_{in} = \tau_{in}^{-1} = \frac{\tau_0}{|\epsilon|^\mu}$  **CRITICAL EXPONENT**  
 $\mu = 1$  IN MFT



At time  $\hat{t}$   $\tau_{in}(\hat{t}) = \hat{t}$  **FREEZE-OUT**

$$\Rightarrow \hat{t} = \frac{\tau_0}{(\hat{t}/\tau_a)^\mu}$$

$$\Rightarrow \hat{t} = (\tau_0 \tau_a^\mu)^{\frac{1}{1+\mu}}$$

$$\epsilon(\hat{t}) = \left(\frac{\tau_0}{\tau_a}\right)^{\frac{1}{1+\mu}}$$

$$\xi = \xi_0 / |\epsilon|^\nu$$

$$\xi^* = \xi(\hat{t}) = \xi_0 \left(\tau_0 / \tau_a\right)^{\frac{\nu}{1+\mu}}$$

$$\text{MFT} \quad \nu = \frac{1}{2} \quad \mu = 1$$

$$\xi^* = \xi_0 \left( \frac{T_0}{T_0} \right)^{\frac{1}{4}}$$

FOR LINE DEFECTS

$$\rho_{\text{DEFECTS}} = (\text{length} / \text{volume}) \text{ defects}$$

$$\sim \frac{1}{(\xi^*)^2}$$

$$\sim \frac{1}{\xi_0^2} \sqrt{\frac{T_0}{T_0}}$$

BUT

- THIS IS AN INTERACTING FIELD THEORY!
- WHAT EFFECT DO THE INTERACTIONS HAVE ON THE DEFECT DENSITY AT FORMATION?

- ▶ INTUITION — EXPECT AN ATTRACTION BETWEEN DOMAINS WITH SMALL PHASE DIFFERENCE RELATIVE TO DOMAINS WITH A LARGE PHASE DIFFERENCE
  - MORE REFINED VERSION OF THE GEODESIC RULE



## SIMULATION

$$\frac{\partial \vec{\phi}}{\partial t} = \vec{\nabla}^2 \vec{\phi} - \frac{\partial V(\vec{\phi})}{\partial \vec{\phi}} + \vec{\zeta}(\vec{r}, t)$$

↑  
GAUSSIAN  
NOISE

$$\langle \vec{\zeta}(\vec{r}, t) \rangle = 0$$

$$\langle \zeta_a(\vec{r}_1, t_1) \zeta_b(\vec{r}_2, t_2) \rangle = 2T \delta_{ab} \cdot \delta(\vec{r}_1 - \vec{r}_2) \delta(t_1 - t_2)$$

LINEAR QUENCH :  $m_f^2 = -m_i^2$

$$T/m_i = \text{CONST.}$$

$$0.001 \leq T/m_i \leq 0.1$$

## PROCEDURE

1. EQUILIBRATE IN DISORDERED PHASE

$$m^2 = m_i^2 \quad \langle \vec{\phi} \rangle = 0 \quad \text{WITH THERMAL NOISE}$$

2. MEASURE LARGEST THERMALLY GENERATED DOMAIN  $V_{\text{MEAN}} \approx 79$

3. QUENCH:  $m^2: m_i^2 \rightarrow -m_i^2$   
 $\langle |\vec{\phi} \cdot \vec{\phi}|^2 \rangle \rightarrow v^2 \neq 0$

4. COMPUTE NO. DEFECTS\*

5. AVERAGE OVER INDEPENDENT SIMULATIONS

6. COMPARE w/ RANDOM DOMAIN FORMATION

\* NUMERICAL IMPLEMENTATION

### DOMAIN DEF<sup>N</sup>

•  $\theta_i, \theta_j \in D$  provided  $|\theta_i - \theta_j| < \theta_c$

•  $\theta_c$   $\left\{ \begin{array}{l} \text{TOO SMALL} \Rightarrow \text{ARTIFICIALLY} \\ \text{LARGE} \neq \text{DOMAINS} \\ \text{TOO LARGE} \Rightarrow \text{DOMAINS LOST} \end{array} \right.$

WE CHOOSE  $\theta_c$  TO BE THE MAXIMUM  
VALUE THAT PRESERVES THE TOPOLOGY  
OF THE ORDER FIELD i.e. REPRODUCES  
DEFECT STRUCTURE OF THE CONTINUUM

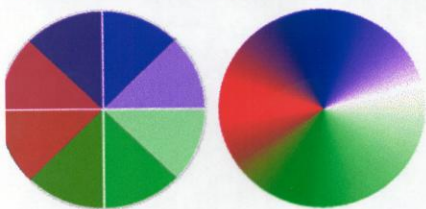
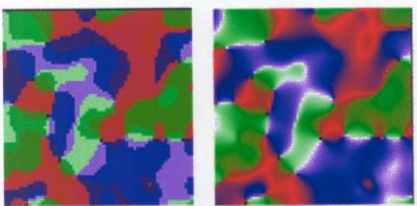
$$S^1 \rightarrow \mathbb{Z}_n = \{e^{2\pi i k/n}, k=0,1,\dots,n-1\}$$

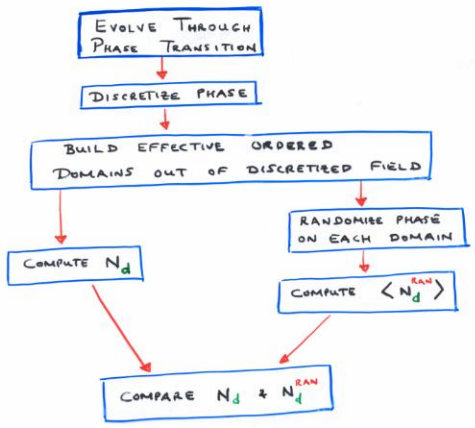
$$\theta_i = \frac{1}{2}(k_i \Delta + (k_i+1)\Delta) \quad \Delta = \frac{2\pi}{n}$$

$$\text{if } \frac{2\pi}{n} k_i < \theta_i < \frac{2\pi}{n} (k_i+1)$$

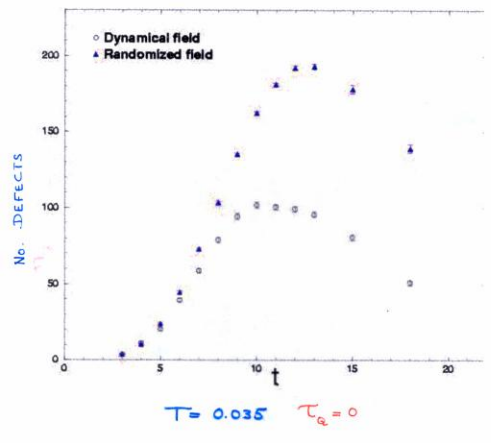


IMPOSE  $\text{VOL}(\text{DOMAIN}) > V_{\text{THERMAL}}$  for  
THERMAL STABILITY



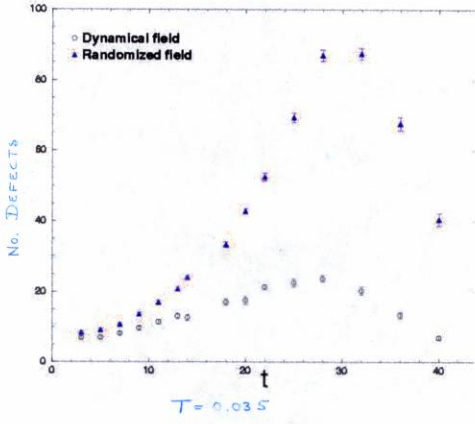


# INSTANTANEOUS QUENCH



FINITE QUENCH RATE

$$T_{QM} = 50$$



## CONCLUSIONS

- NAIVE CORRELATION LENGTH  $\equiv$  MAXIMAL DOMAIN SIZE AT FREEZE OUT  
DOES NOT GIVE THE CORRECT TOPOLOGICAL DEFECT DENSITY FOR FINITE QUENCH RATES
- PHASE ALIGNMENT VIA DOMAIN INTERACTION REDUCES WANDERING ON THE GROUND STATE MANIFOLD  $\mathcal{M}$  i.e. PHASES/ORDER FIELDS ARE NOT RANDOM AT THE TIME OF DEFECT FORMATION

## FUTURE

- DETERMINE  $T_q$  WHICH GIVES MAXIMAL SUPPRESSION OF DEFECT DENSITY
- RELATIVISTIC DYNAMICS (2ND-ORDER IN TIME)
- DYNAMICAL SPACETIME