

Cindy Regal

Condensed Matter Summer School, 2018

Day 1: Quantum optomechanics

Day 2: Quantum transduction

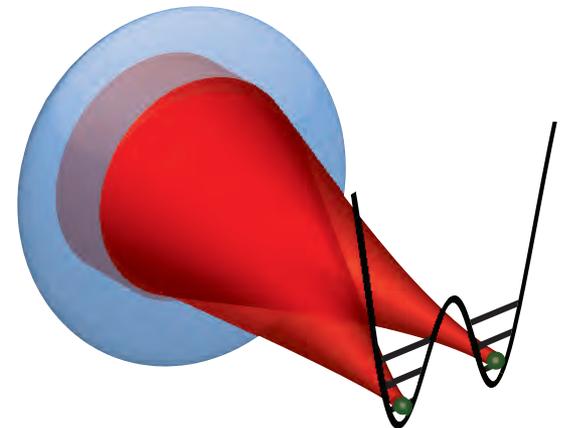
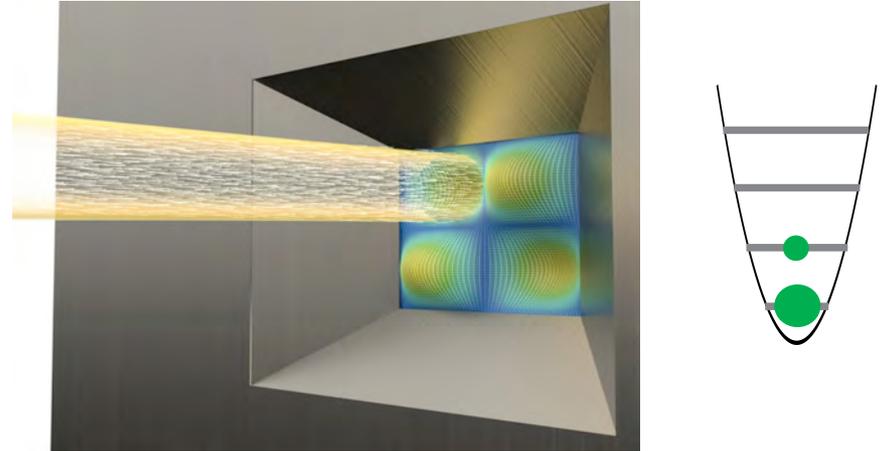
Day 3: Ultracold atoms from a qubit perspective



Day 1: Quantum optomechanics

Day 2: Quantum transduction

Day 3: Ultracold atoms from a qubit perspective



Day 1: Quantum optomechanics – quantum limits to continuous displacement detection

Day 2: Quantum transduction – conversion from microwave (superconducting qubits) to optical photons (transmission domain)

Machinery is that of weak nonlinearity / gaussian states

Useful to understanding from perspective of quantum metrology, transducers

Day 3: Ultracold atoms from a qubit perspective – interfering and entangling bosonic atoms

Single atom ‘sources’

Overview of field of control of individual neutral atoms

Some examples of creating Bell states

Quantum optomechanics outline

Basic interaction and cooling of moving mirror in cavity / interferometer
– example of more general machinery

Example optomechanics experiments

Continuous displacement detection and squeezing – an old problem

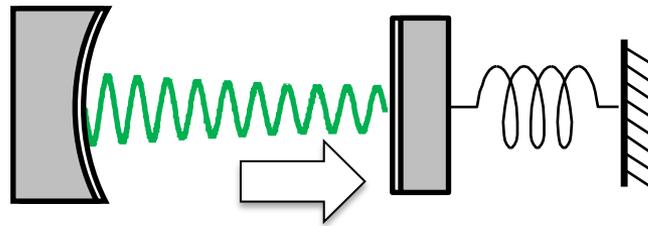
Quantum noise and measurement and amplifiers

Some of our experiments

Will lead into tomorrow – can we optically read out the force of a single microwave photon? – modern problem of quantum transduction from superconducting qubits

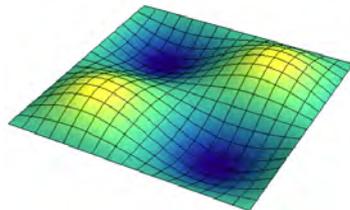
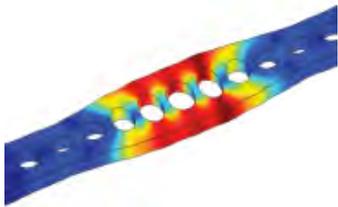
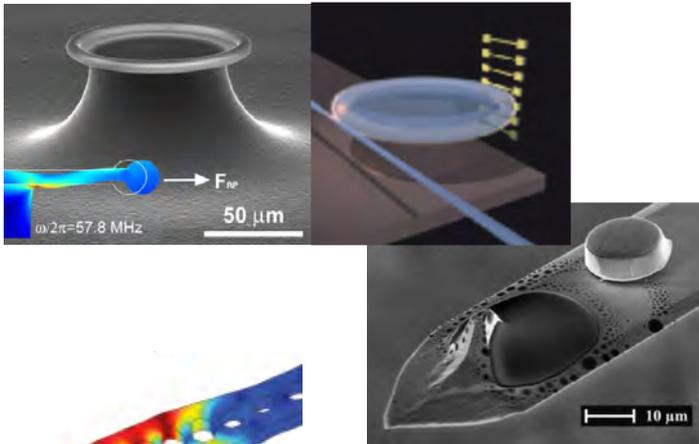
Light controls micromechanical motion

Micromechanical motion controls light

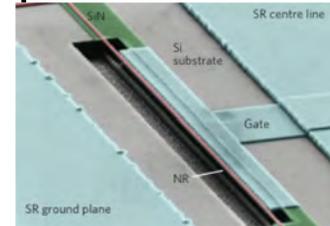
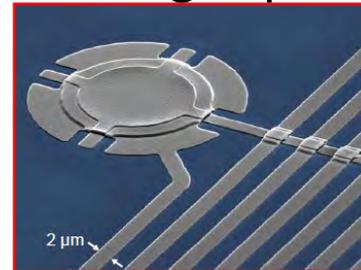


Radiation pressure

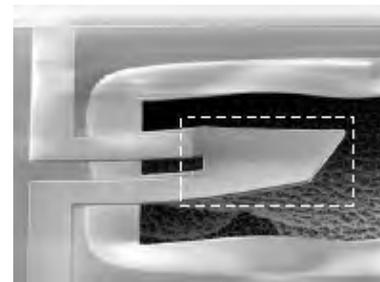
Effectively moving mirrors



Moving capacitor plate



Piezoelectrics



Optomechanical interaction

$$\hat{F} = \hbar G \hat{n} \quad \hat{n} = \hat{a}^\dagger \hat{a}$$

$\hookrightarrow \frac{d\omega_c}{dx}$

$\hat{a} \rightarrow$ cavity mode

$$\hat{H}_{\text{int}} = \hbar G \hat{a}^\dagger \hat{a} x_{\text{zp}} \hat{x}$$
$$= \hbar g \hat{a}^\dagger \hat{a} \hat{x}$$

$\hat{b} \rightarrow$ mechanical mode

$$\hat{H} = \hbar \omega_c \hat{a}^\dagger \hat{a} + \hbar \omega_m \hat{b}^\dagger \hat{b} + \hbar g \hat{a}^\dagger \hat{a} (\hat{b}^\dagger + \hat{b})$$

$\hookrightarrow \frac{d\omega_c}{dx} x_{\text{zp}}$

$$\hat{a} \rightarrow \alpha + \zeta \hat{a}$$

\propto stemming from large $\cup \alpha_{\text{in}}$

Optomechanical interaction

$$\hat{H}_{\text{int}} = \hbar g (\alpha + \hat{s}_a)^\dagger (\alpha + \hat{s}_a) (\hat{b} + \hat{b}^\dagger)$$

$$= \hbar g |\alpha|^2 (\hat{b} + \hat{b}^\dagger) + \hbar g (\alpha^* \hat{s}_a + \alpha \hat{s}_a^\dagger) (\hat{b} + \hat{b}^\dagger)$$

avg vp

$$+ \hbar g \hat{s}_a \hat{s}_a^\dagger (\hat{b} + \hat{b}^\dagger)$$

smaller by α

α real valued $\alpha = \sqrt{N}$

$$\hat{H}_{\text{int}}^{(\text{lin})} = \hbar g \alpha (\hat{s}_a + \hat{s}_a^\dagger) (\hat{b} + \hat{b}^\dagger)$$

$$\Delta = \omega_L - \omega_c$$

Red detuned $\Delta = -\omega_m$

$$\hat{s}_a^\dagger \hat{b} + \hat{s}_a \hat{b}^\dagger$$

Beamsplitter

On resonance $\Delta = 0$

$$(\hat{s}_a^\dagger + \hat{s}_a) \hat{x}$$

Interferometer

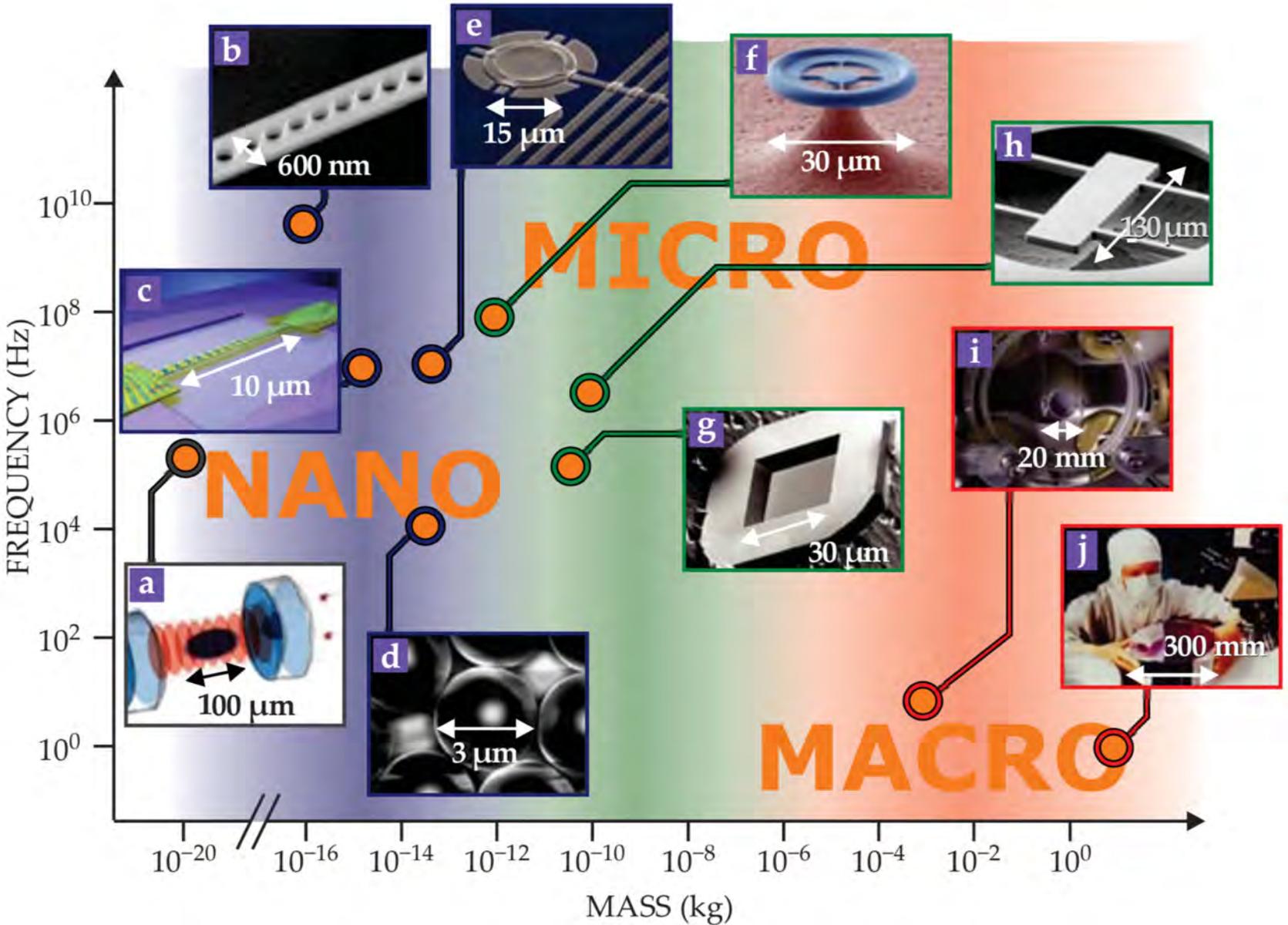
Blue detuned $\Delta = \omega_m$

$$\hat{b} \hat{s}_a + \hat{b}^\dagger \hat{s}_a^\dagger$$

Two-mode

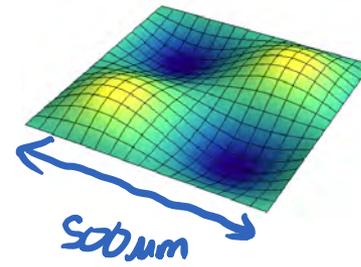
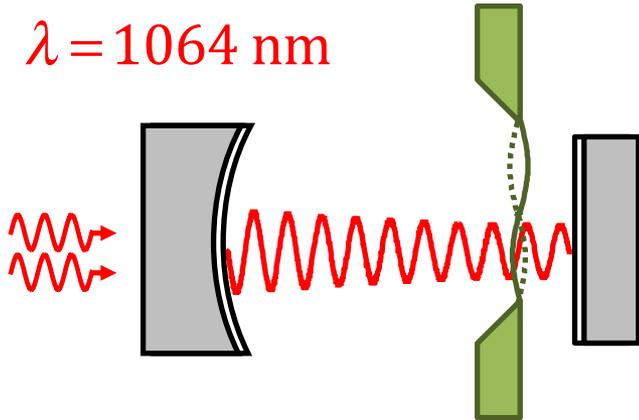
squeezing

Range of size scales



Our optomechanical device

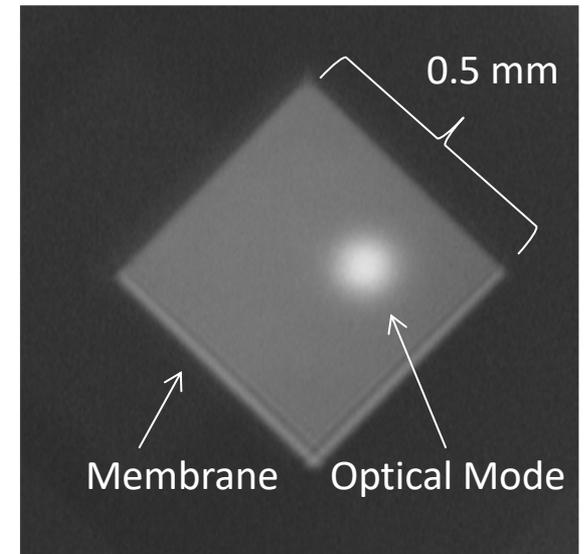
$$\lambda = 1064 \text{ nm}$$



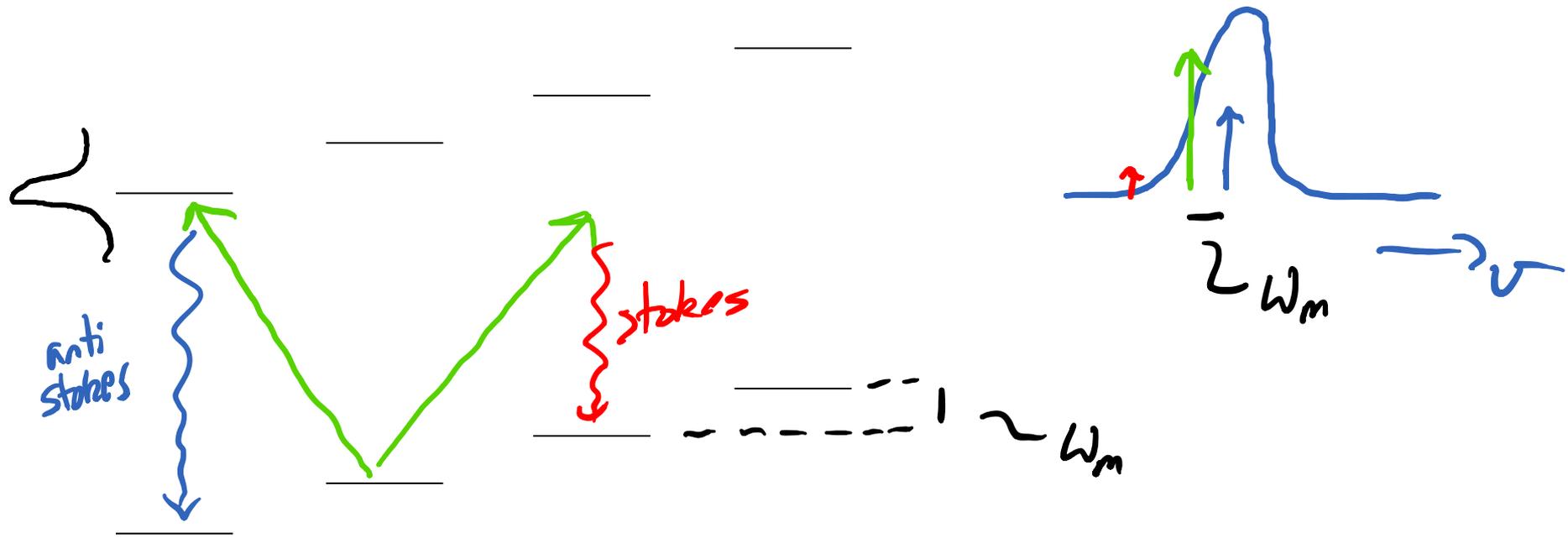
500 μm
50 nm thick

$$m \sim 10 \text{ ng}$$
$$\nu_m = 1 \text{ MHz}$$

$$x_{zp} = \sqrt{\frac{\hbar}{2m\omega_m}}$$
$$\sim 1 \text{ fm}$$



Resolved sideband cooling



$$\propto \frac{1}{1 + (4\omega_m/k)^2} \times (n_m + 1)$$

$$\propto n_m$$

$$\frac{S_{FF}(\omega_m)}{S_{FF}(-\omega_m)} = \frac{n_m + 1}{n_m} = e^{+\hbar\omega/k_B T}$$

Reaching 'quantum' limit

Optical measurement rate /
Photon-phonon exchange
rate with propagating field

LARGE

$$\Gamma_{\text{opt}} = \frac{4\alpha^2 g^2}{K}$$

$$\alpha^2 = N$$

$$\left\{ \begin{array}{l} \text{Lots of light} \\ g = \chi_{zp} \frac{dW_c}{dx} \end{array} \right.$$

Thermalization rate with
mechanical environment

slow

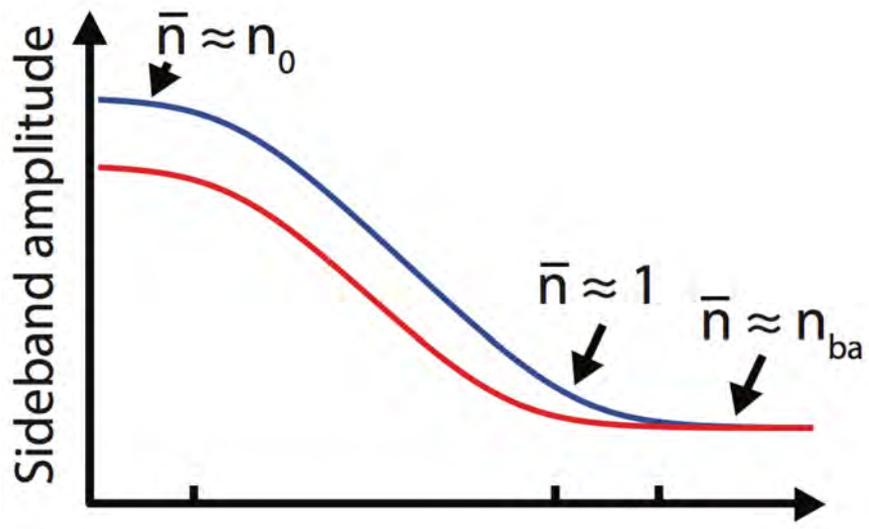
$$\Gamma_{\text{th}} = n_{\text{th}} \Gamma_m$$

$$\Gamma_m = \frac{W_m}{Q_m}$$

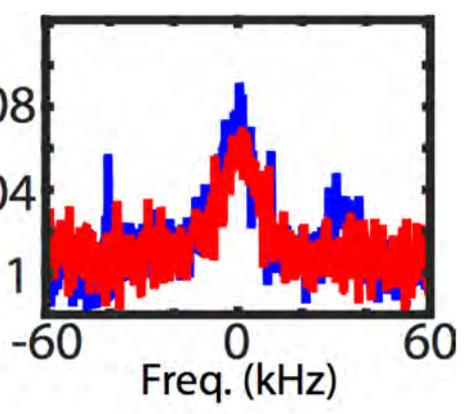
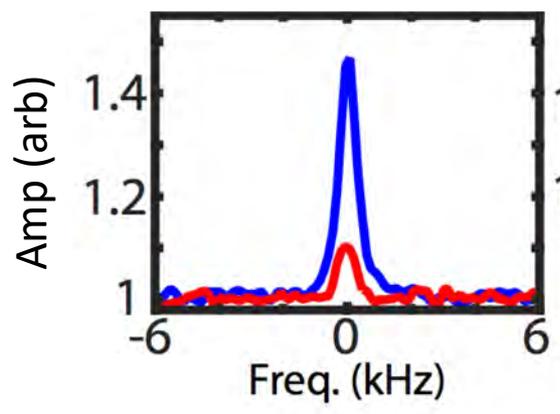
$$n_{\text{th}} = \frac{k_b T}{\hbar W_m}$$

$$\left\{ \begin{array}{l} \text{High } Q_m \\ \text{Cold Bath} \end{array} \right.$$

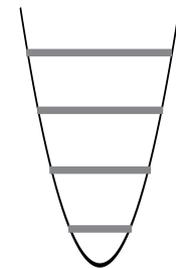
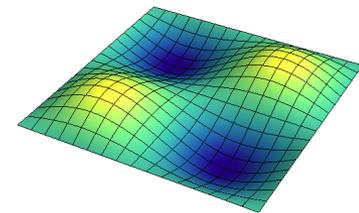
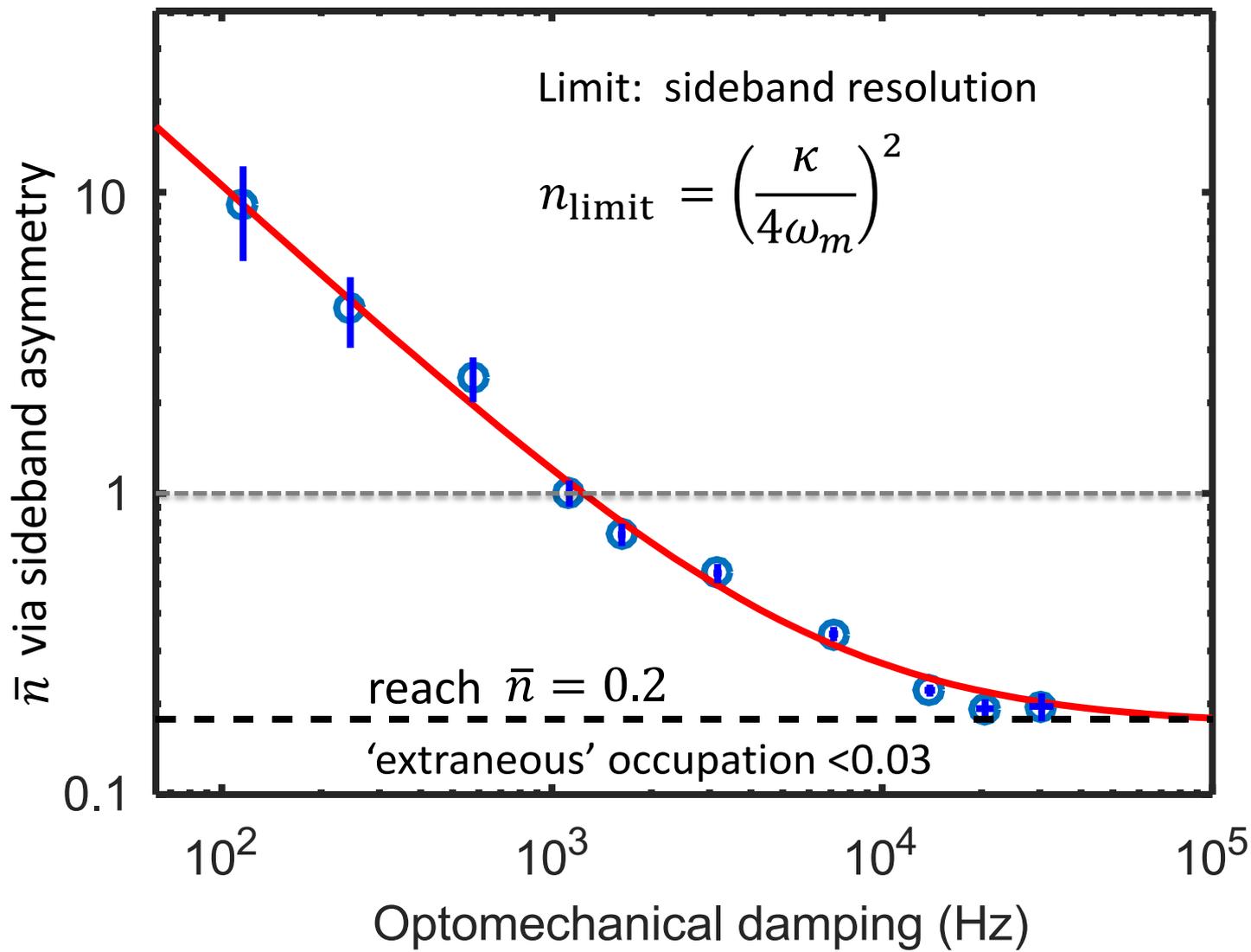
Quantum backaction limit of cooling



Γ_{opt}



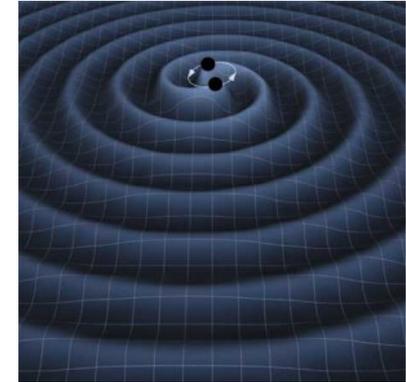
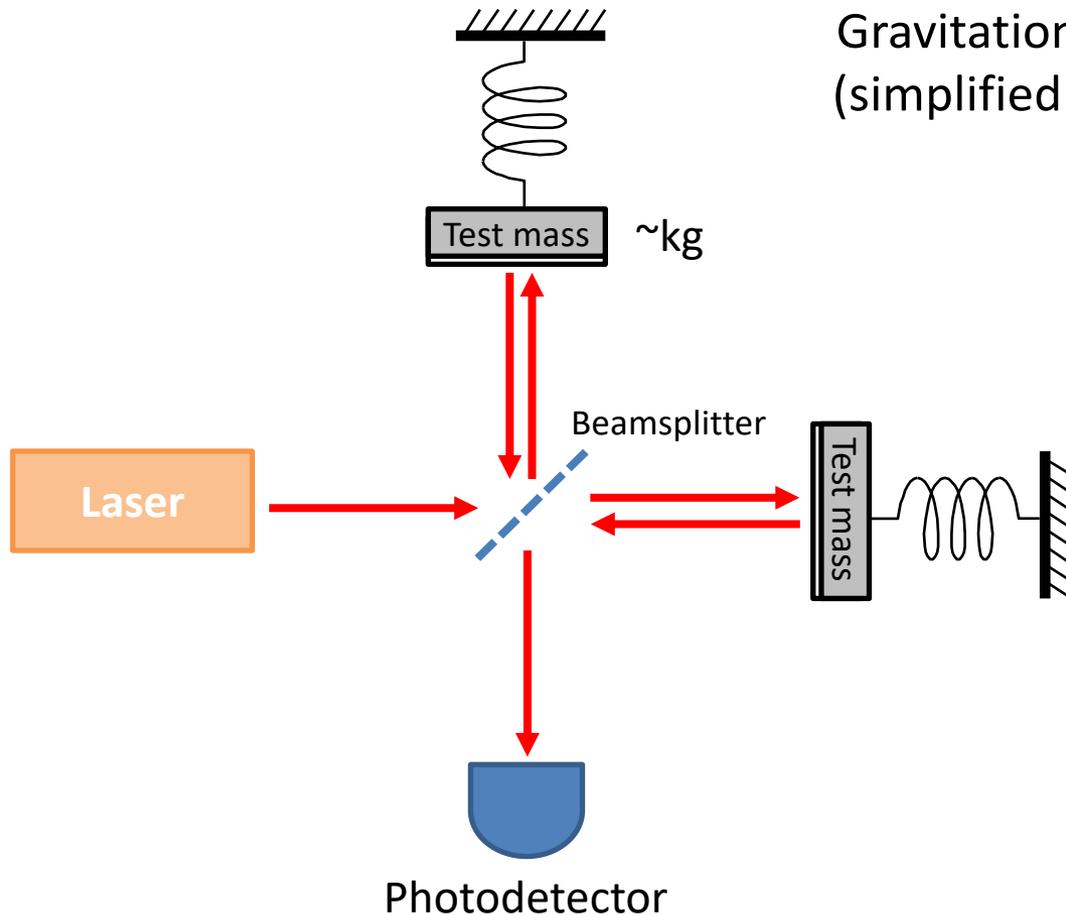
Ground-state optomechanical cooling



Data here: R. W. Peterson *et al.*, PRL (2016)

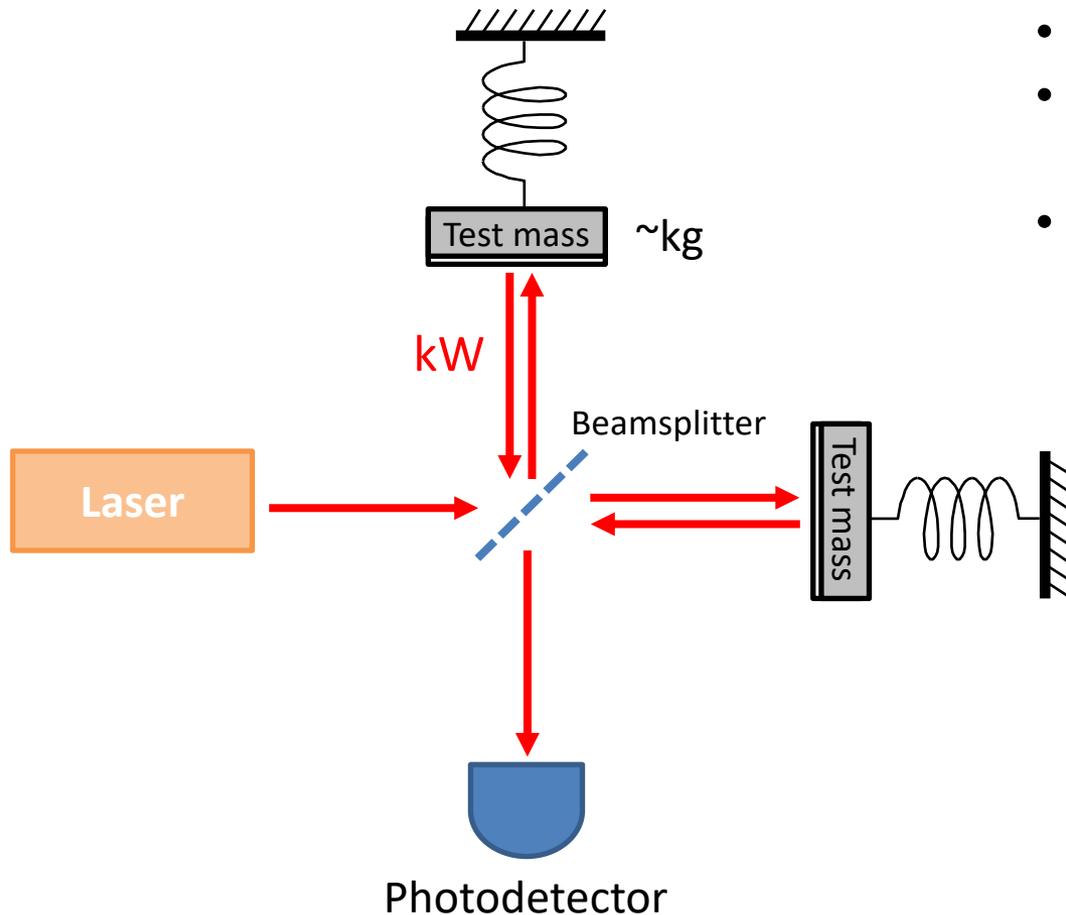
Ground state cooling: Teufel *et al.*, Nature (2011), Chan *et al.*, Nature (2012), Khalili *et al.*, PRA (2012)

Gravitational wave detection



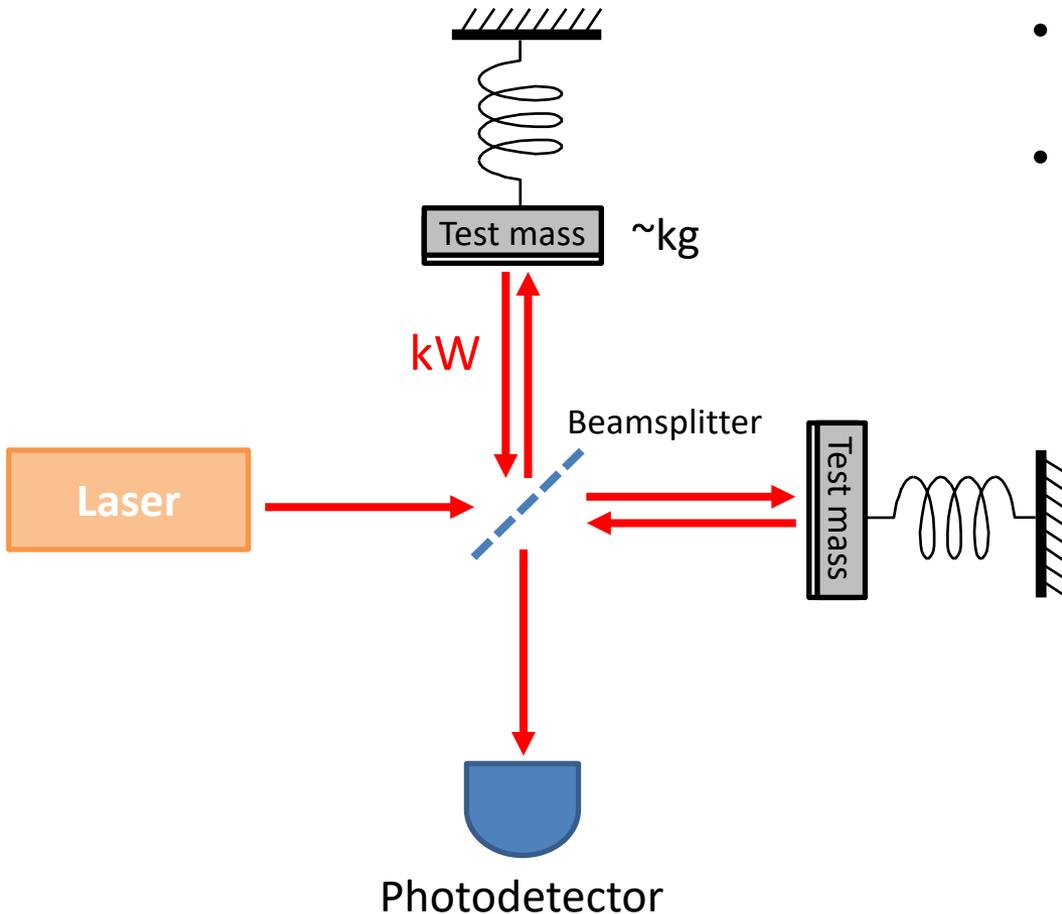
- Sensitive optical interferometer
- Aims to detect $\sim 10^{-18}$ m

In limit of strong probing



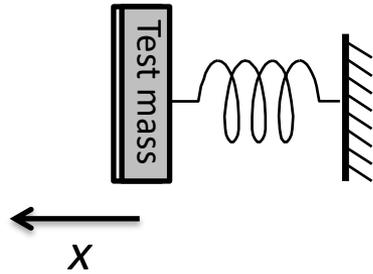
- No longer immutable structure
- Light pressure on mirrors important – radiation pressure
- Vladimir Braginsky, 1970's – instabilities

In limit of strong probing



- Effect on signal to noise?
- Will radiation forces combined with fluctuations of light obscure the motion?

Continuous measurement: Round 1



- *Back of the envelope* calculation (Heisenberg microscope argument)
- Consider free mass limit - time small compared to oscillation period
- To measure passing wave must measure more than once...

Measure to Δx_{meas} @ time t

Momentum uncertain to $\Delta p_{\text{perturb}} \geq \frac{\hbar}{2 \Delta x_{\text{meas}}}$

At a later time....

$$\Delta x(t') = \Delta x(t) + \frac{\hbar(t' - t)}{2m \Delta x(t)} \rightarrow \Delta x_{\text{SQL}} = \sqrt{\frac{2\hbar(t' - t)}{m}}$$

Source of backaction: Radiation pressure SN

Quantum-Mechanical Radiation-Pressure Fluctuations in an Interferometer

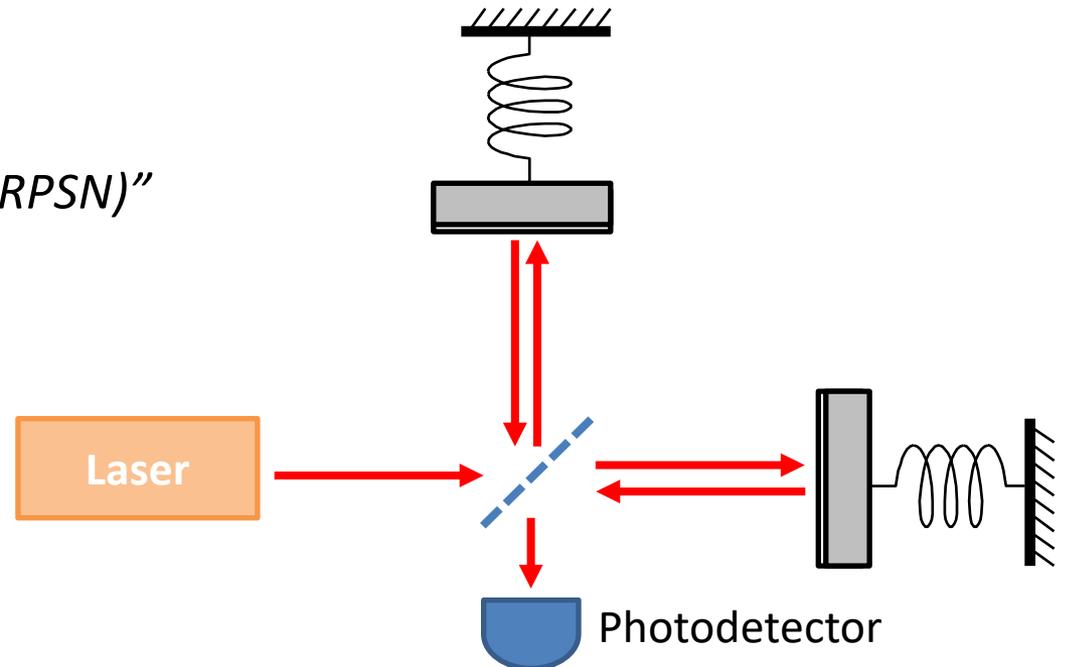
Carlton M. Caves

W. K. Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, California 91125
(Received 29 January 1980)

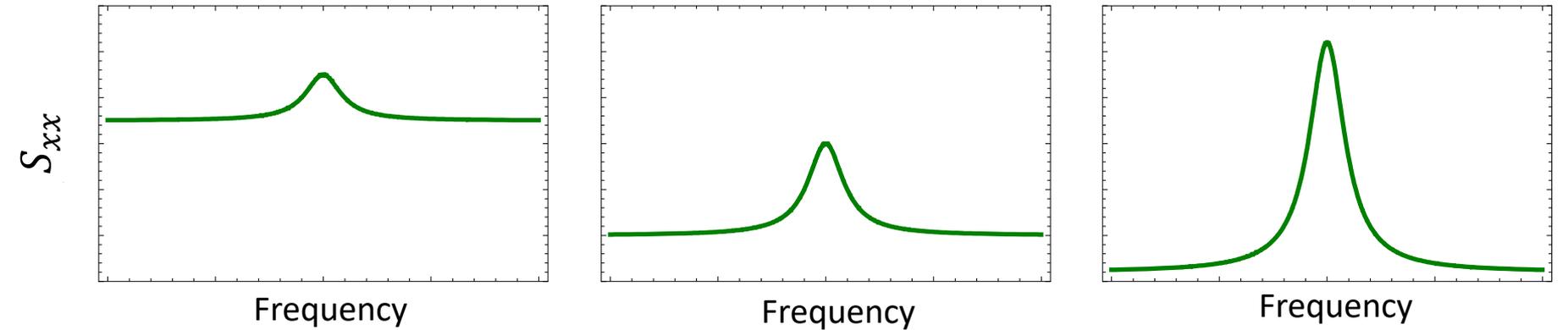
The interferometers now being developed to detect gravitational waves work by measuring small changes in the positions of free masses. There has been a controversy whether quantum-mechanical radiation-pressure fluctuations disturb this measurement. This Letter resolves the controversy: They do.

Our terminology:

“Radiation pressure shot noise (RPSN)”



Standard quantum limit in continuous detection



$\rightarrow \omega$

S_{xx}
units m^2/Hz

\Rightarrow

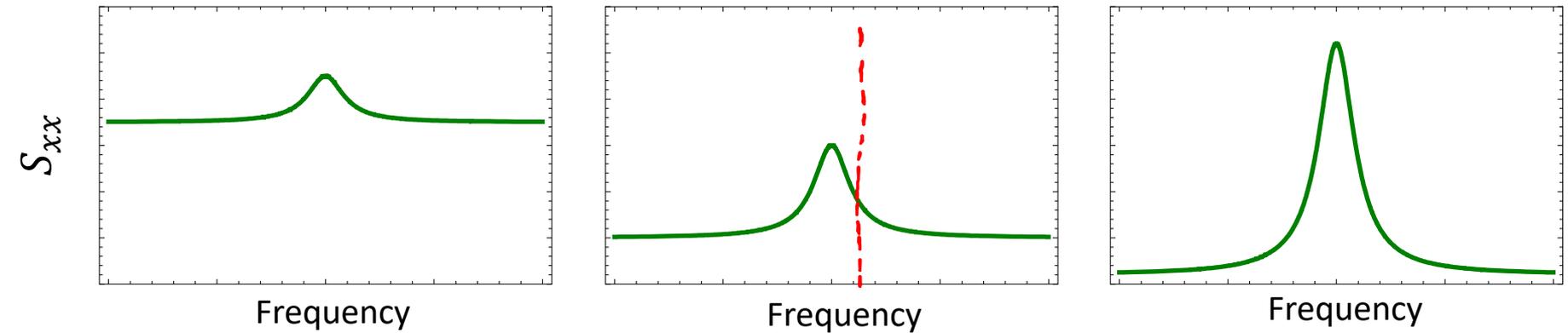
Imprecision S_{II}

Backaction force S_{FF}

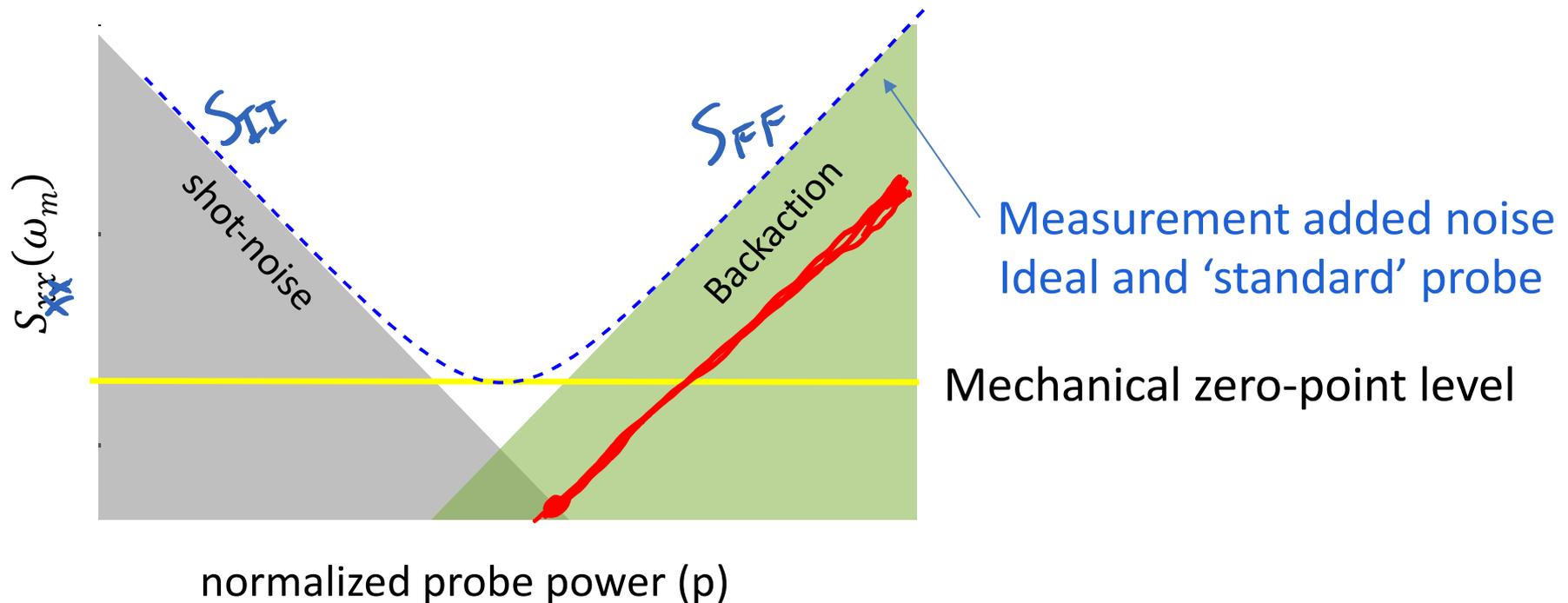
$\rightarrow |\chi_m|^2 S_{FF}$

$$\chi_m = \frac{1}{m(\omega^2 - \omega_m^2) + im\Gamma_m\omega}$$

Standard quantum limit in continuous detection



SQL on mechanical resonance



Added noise for two-quadrature measurements

$$X_1 = \frac{1}{2}(a_s + a_s^\dagger)$$

$$X_2 = \frac{1}{2i}(a_s - a_s^\dagger)$$

$$[X_1, X_2] = \frac{i}{2}$$

uncertainty product

$$\langle \Delta X_1^2 \rangle \langle \Delta X_2^2 \rangle \geq \frac{1}{16}$$

Added noise $A_n = 1/2$

Amplify

$$Y_1 = \sqrt{G_1} X_1 + F_1$$

$$Y_2 = \sqrt{G_2} X_2 + F_2$$

$$\text{Also need } [Y_1, Y_2] = \frac{i}{2}$$

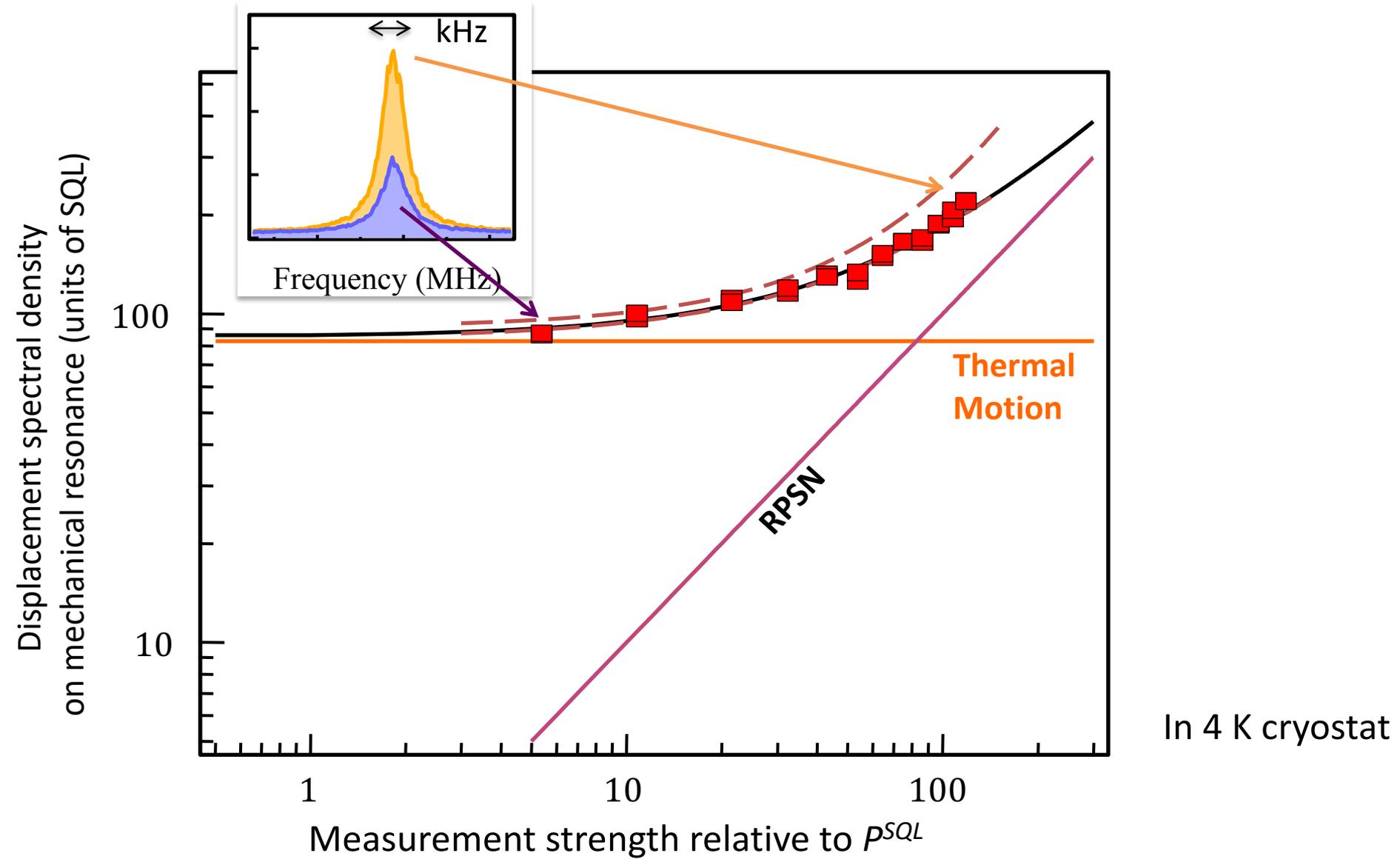
Ends up implying

$$\frac{\langle \Delta Y_1^2 \rangle}{G_1} \frac{\langle \Delta Y_2^2 \rangle}{G_2} \geq \frac{1}{16} \left(2 - \frac{1}{\sqrt{G_1} \sqrt{G_2}} \right)^2$$

large G

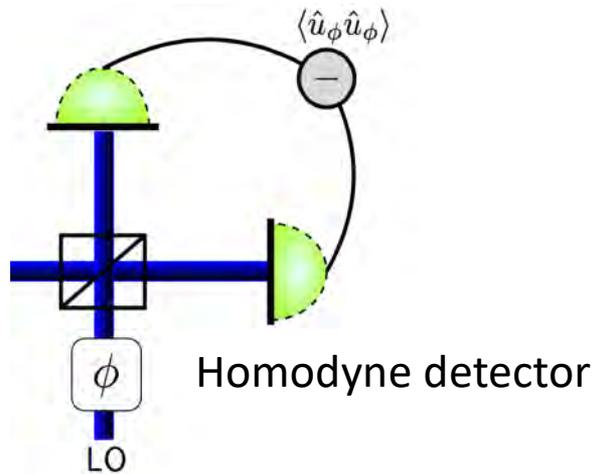
Experiment: Observation of RPSN

RPSN: Radiation Pressure Shot Noise

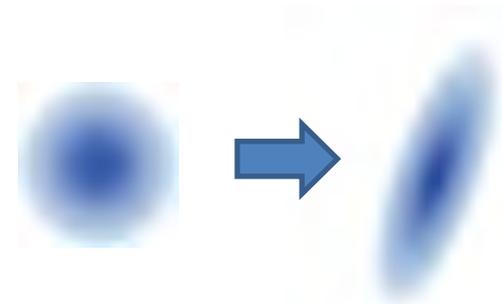
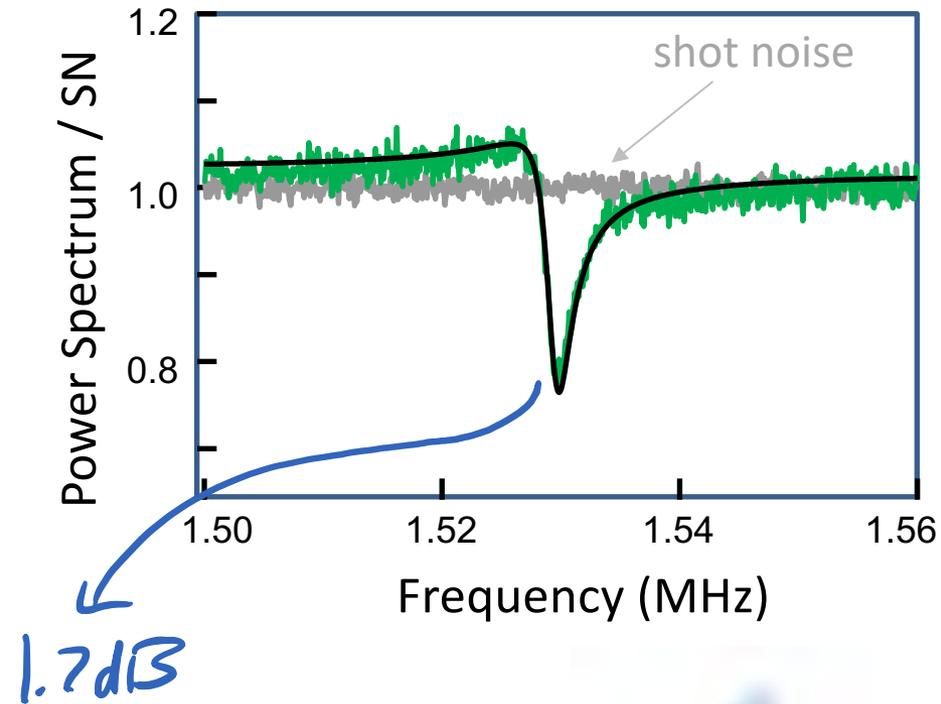


Squeezing of light: Correlations in quantum noise

Radiation pressure shot noise (amplitude of light) drives mechanics, writes back onto cavity (phase of light)



Ponderomotive Squeezing



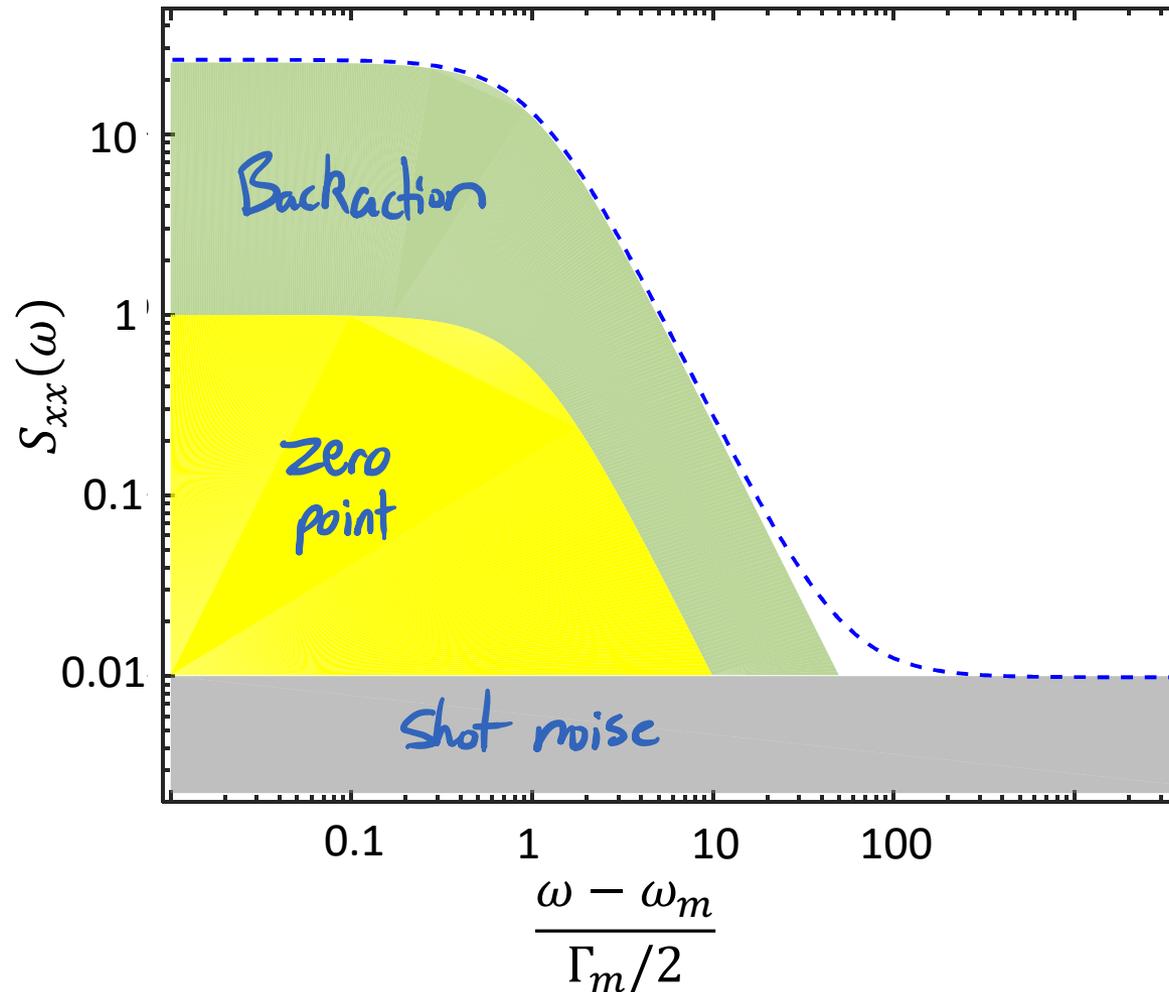
Optomechanical squeezing observed with:

Cold gases: D. W. Brooks...D. M. Stamper-Kurn, Nature (2012)

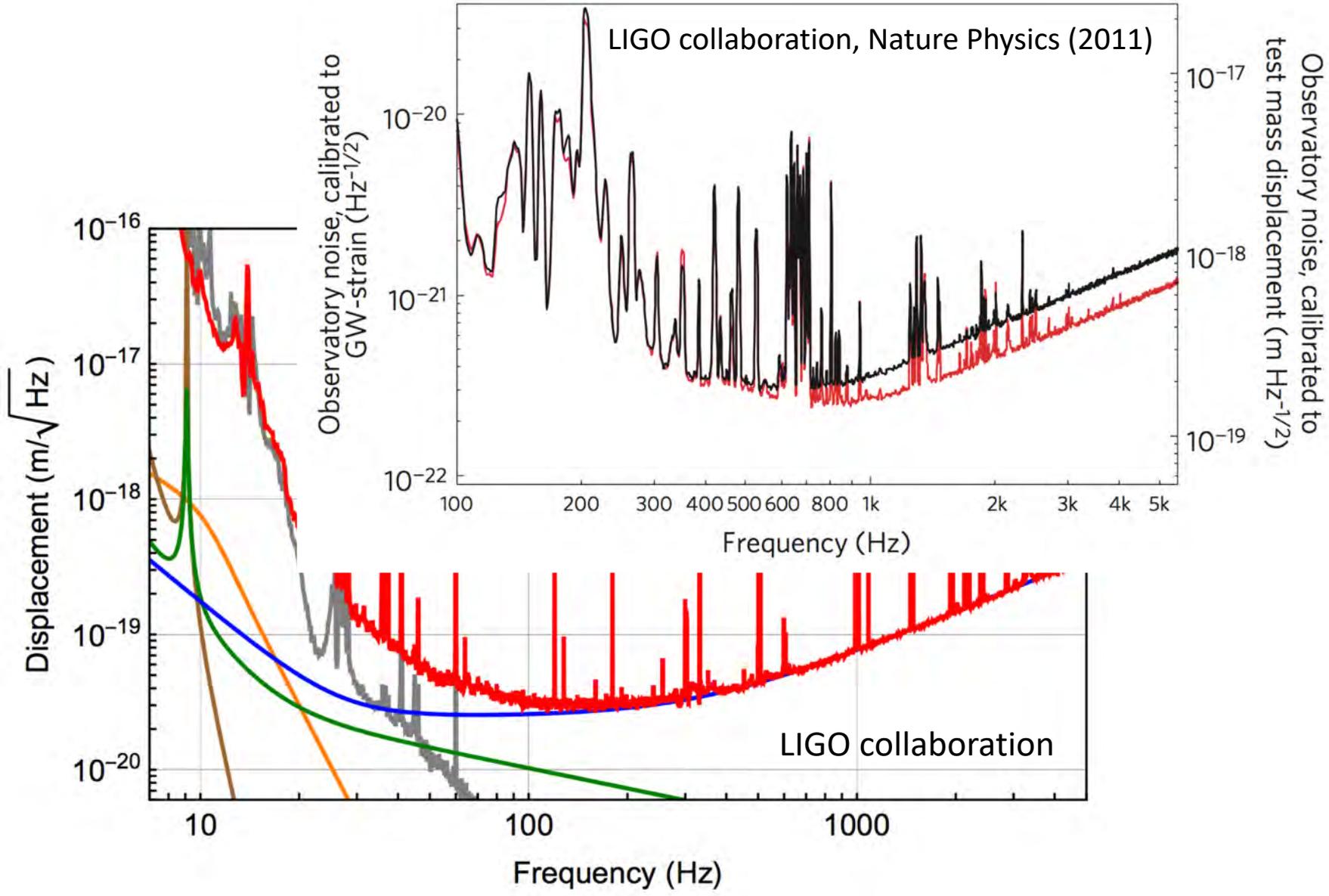
Nanomechanical devices: A.-H. Safavi-Naeini...O. Painter, Nature (2013); T. P. Purdy *et. al.*, PRX (2013)

Illustrated guide to broadband detection

Frequency dependence for backaction-limited probe

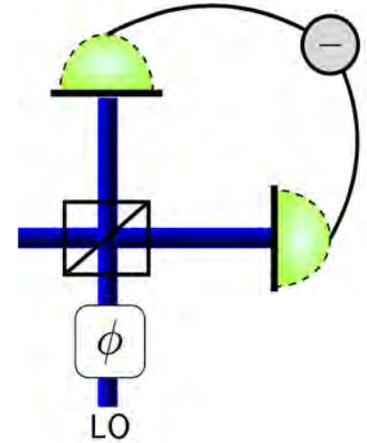
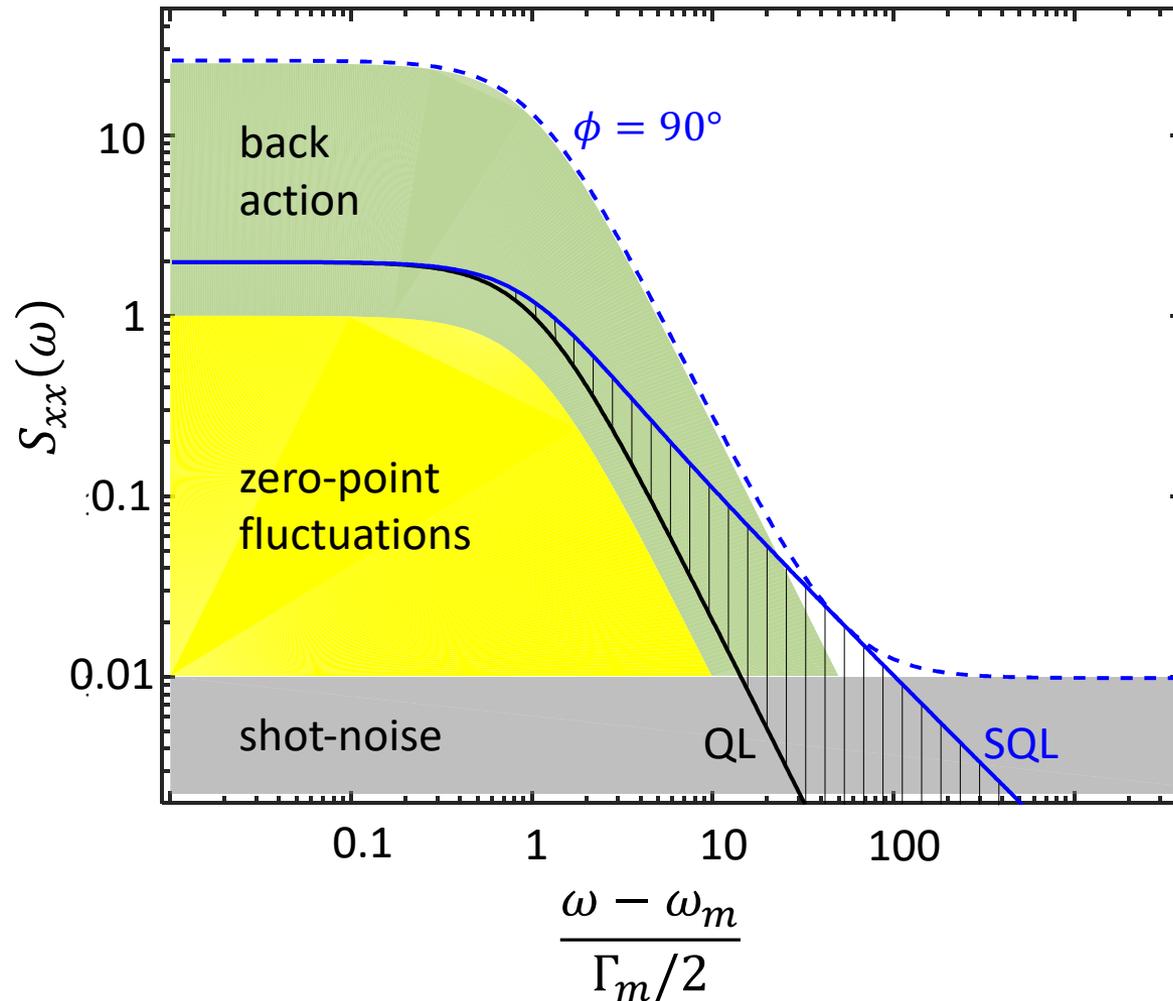


LIGO measurement context



Illustrated guide to broadband detection

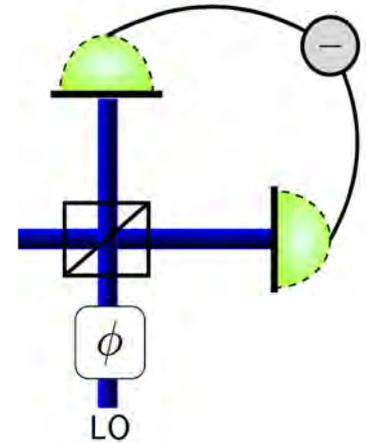
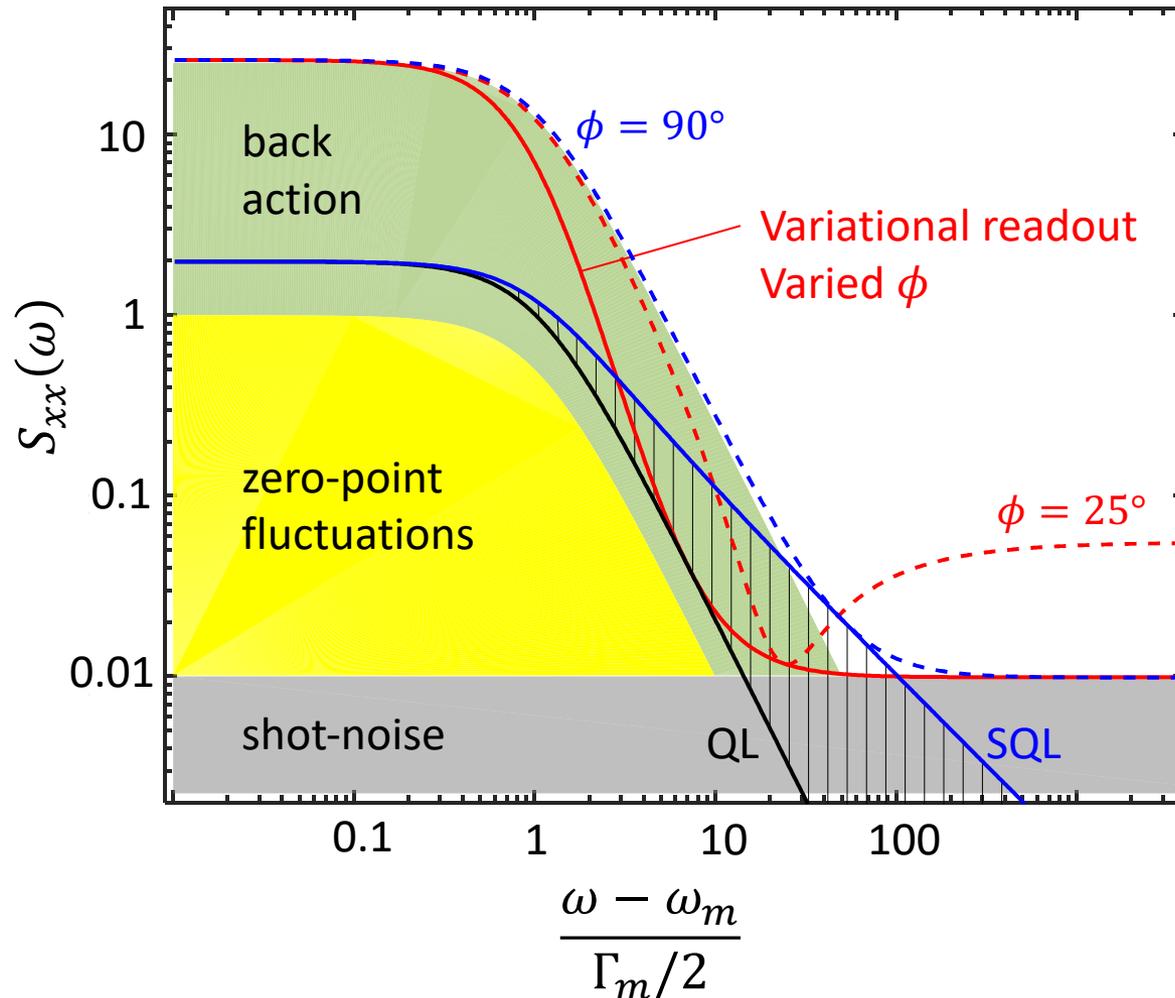
$$X_\phi = X_{AM} \cos\phi + X_{PM} \sin\phi$$



QL – quantum limit for two mechanical quadrature measurement

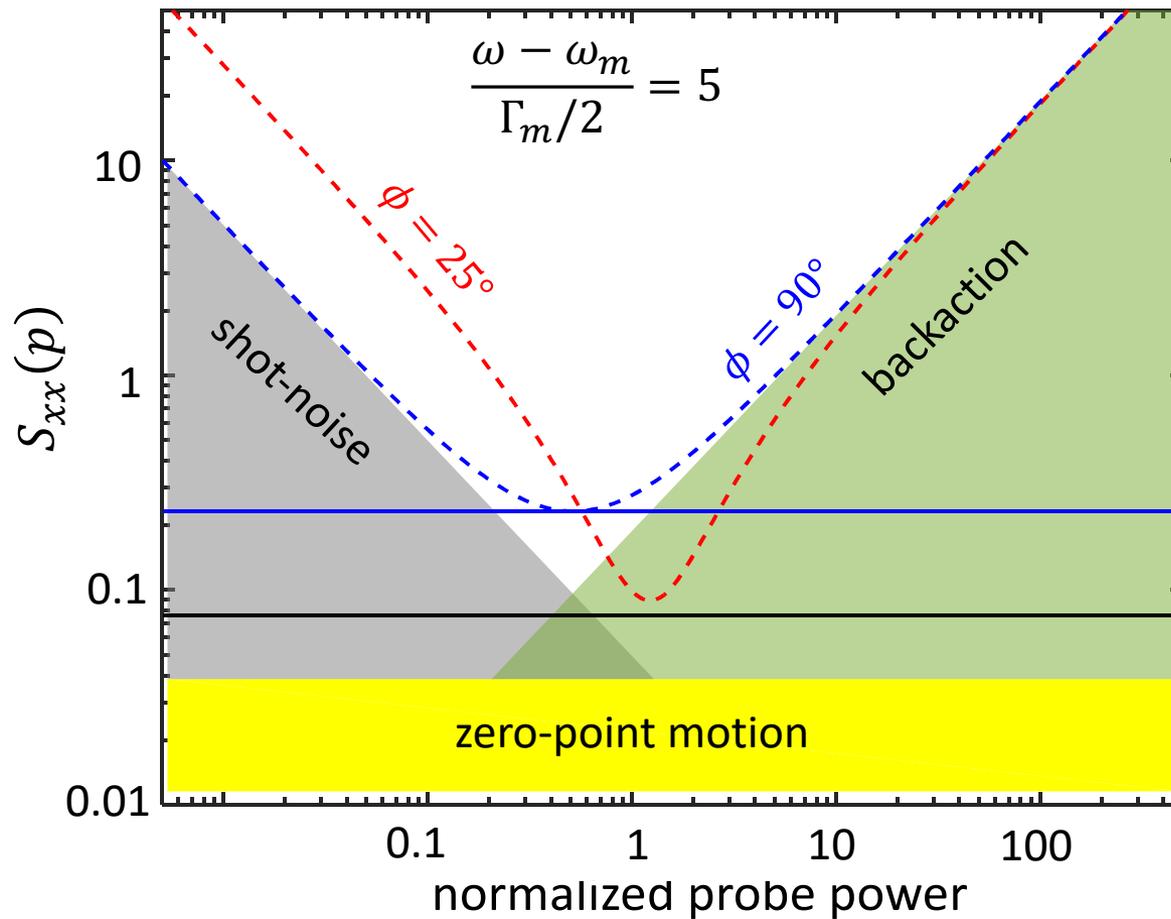
Using ponderomotive squeezing

$$X_\phi = X_{AM} \cos\phi + X_{PM} \sin\phi$$



As a function of probe power

$$X_\phi = X_{AM} \cos\phi + X_{PM} \sin\phi$$

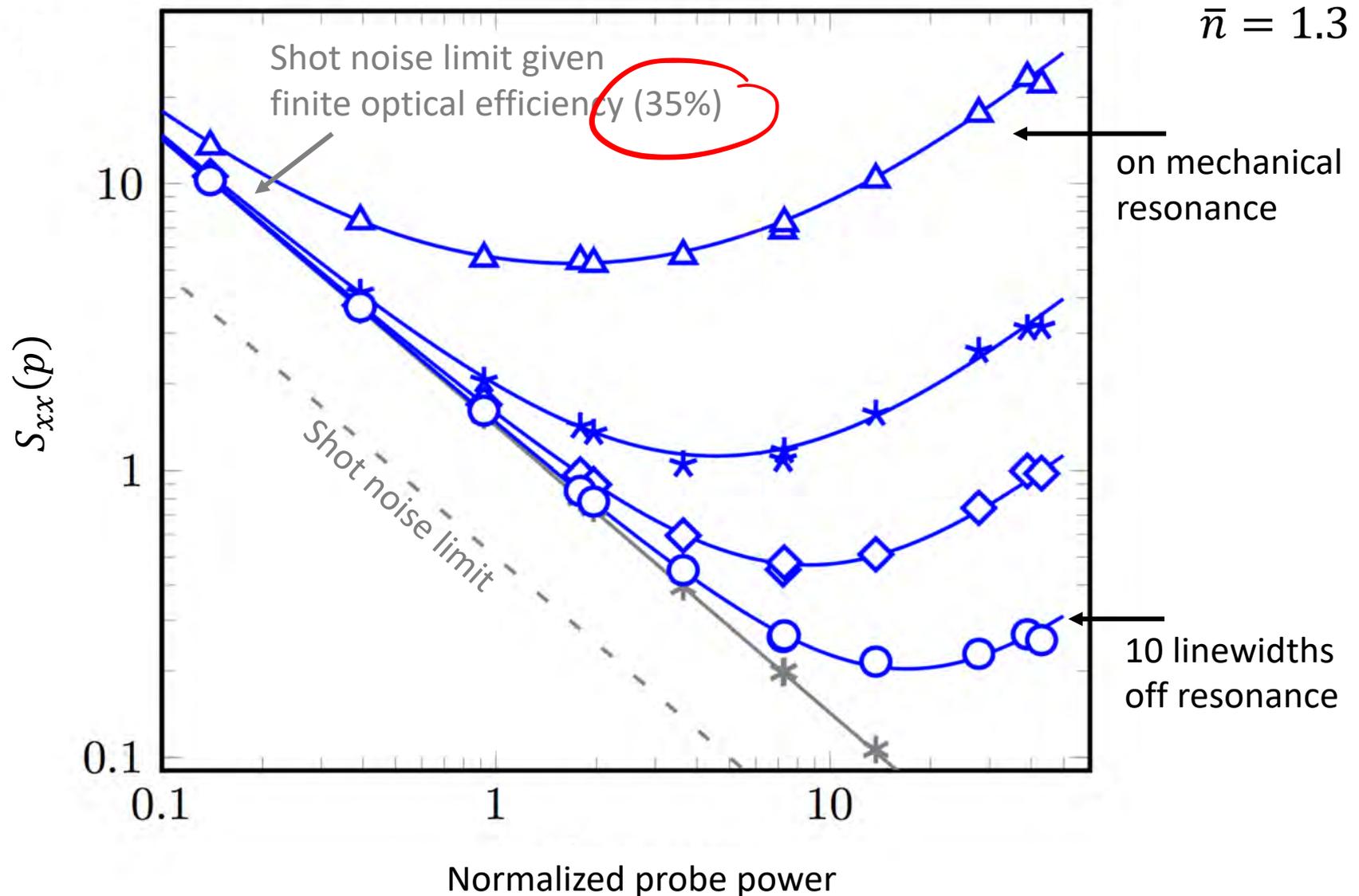


$$S_{xx}(\omega) = S_m(\omega) + S_{II} + |\chi_m(\omega)|^2 S_{FF} \\ + 2 \operatorname{Re} [\chi_m(\omega) S_{IF}]$$

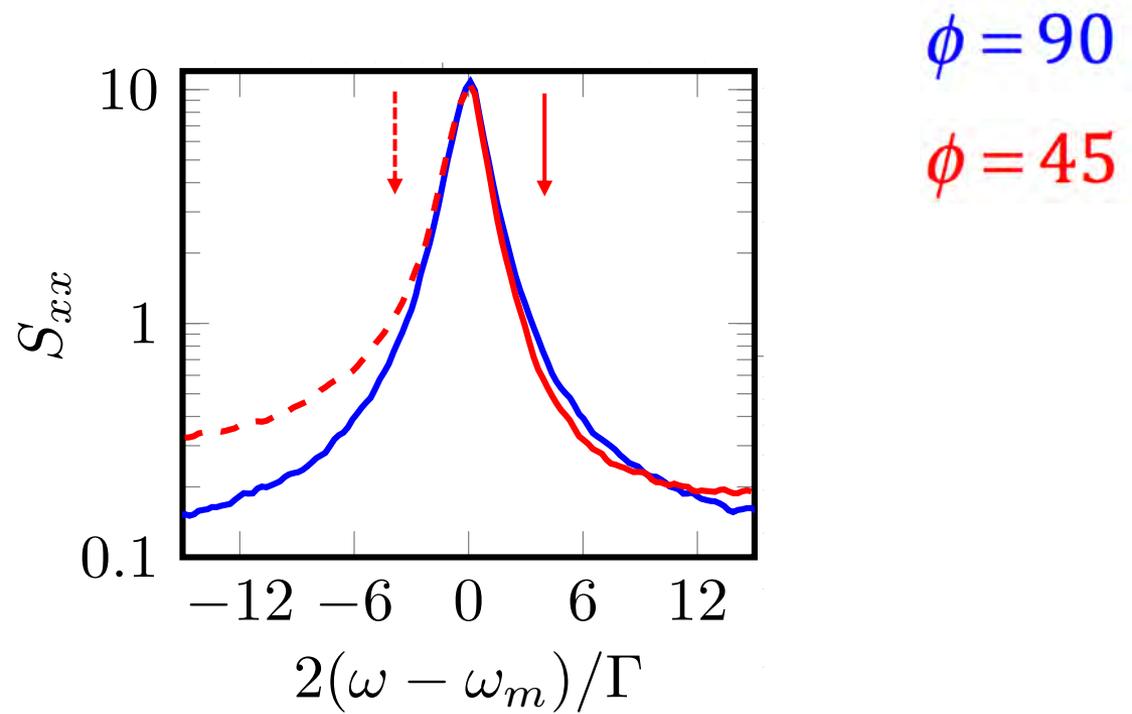
$$\chi_m(\omega) = \frac{1}{m(\omega^2 - \omega_m^2) + i \Gamma_{mm} \omega}$$

SQL and off-resonant SQL measurements

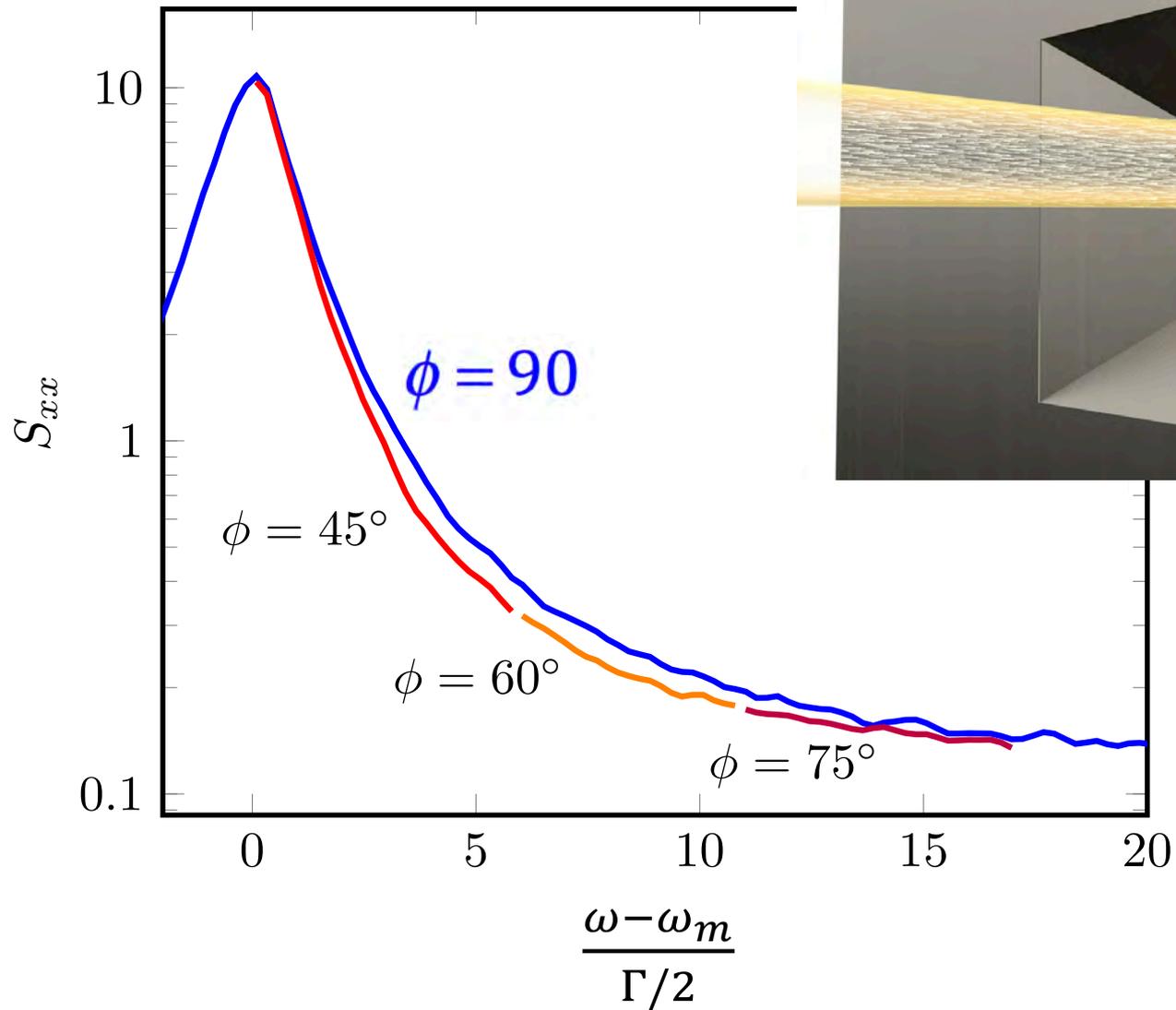
Probe damped mechanics with on-cavity-resonance probe



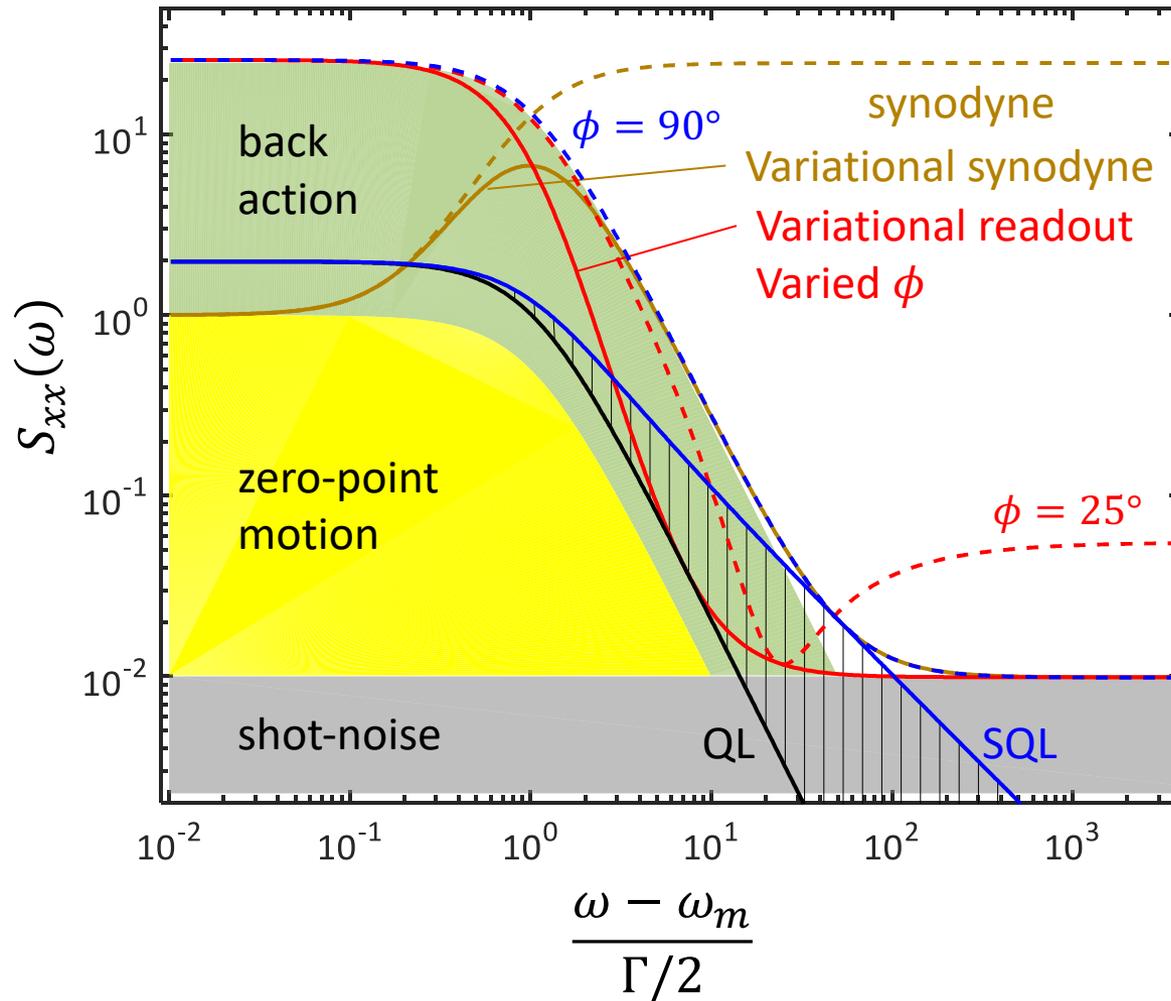
SQL and off-resonant SQL measurements



Idea of variational readout

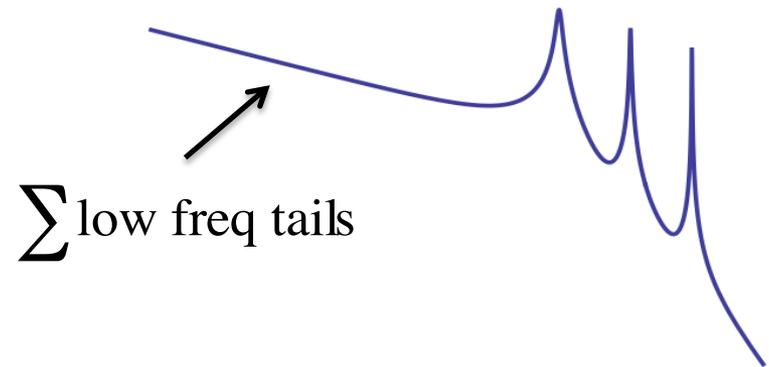
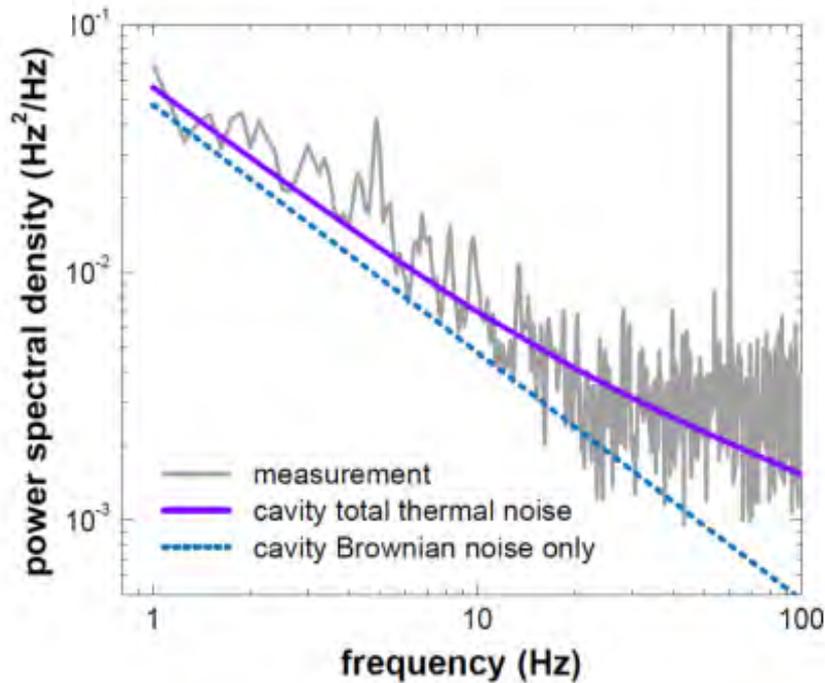


Illustrated guide to broadband detection



Thermal noise – ubiquitous problem in metrology

State of the art reference cavities

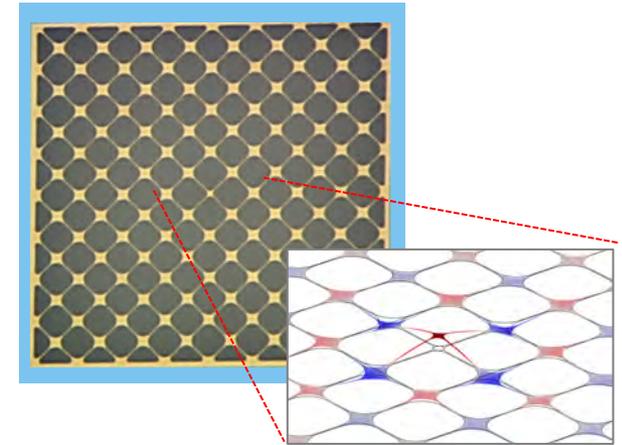
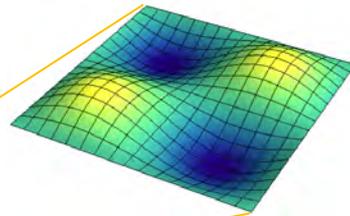
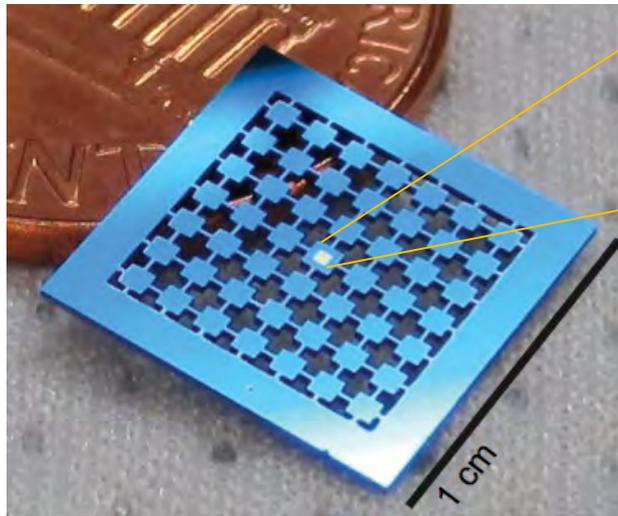


- Tail grows with material loss (smaller Q)
- Crystalline materials are desired
- We live in a forest of modes at higher frequency

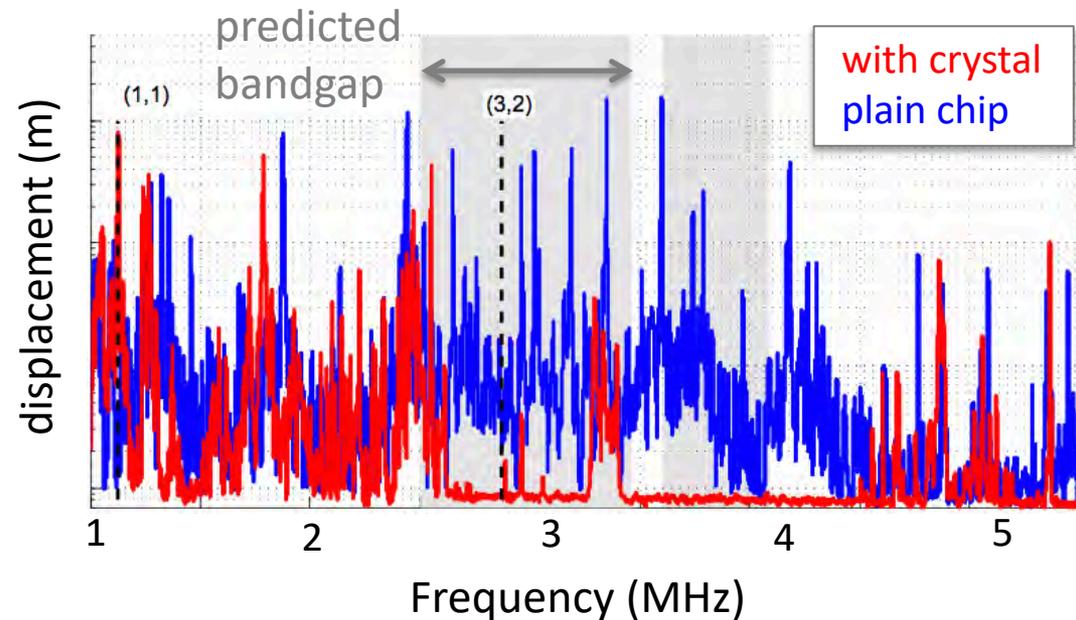
Phononic crystals

Period structures: Phononic crystal

- Control mode structure
- Reduce acoustic energy at lossy boundary



Displacement on membrane

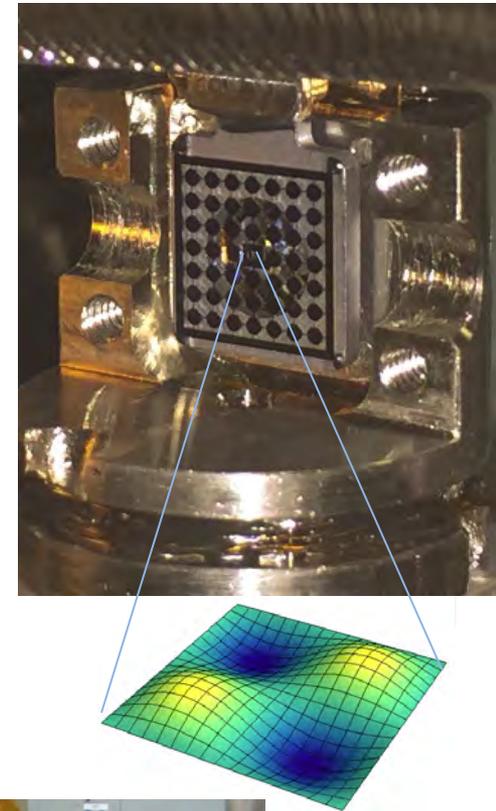
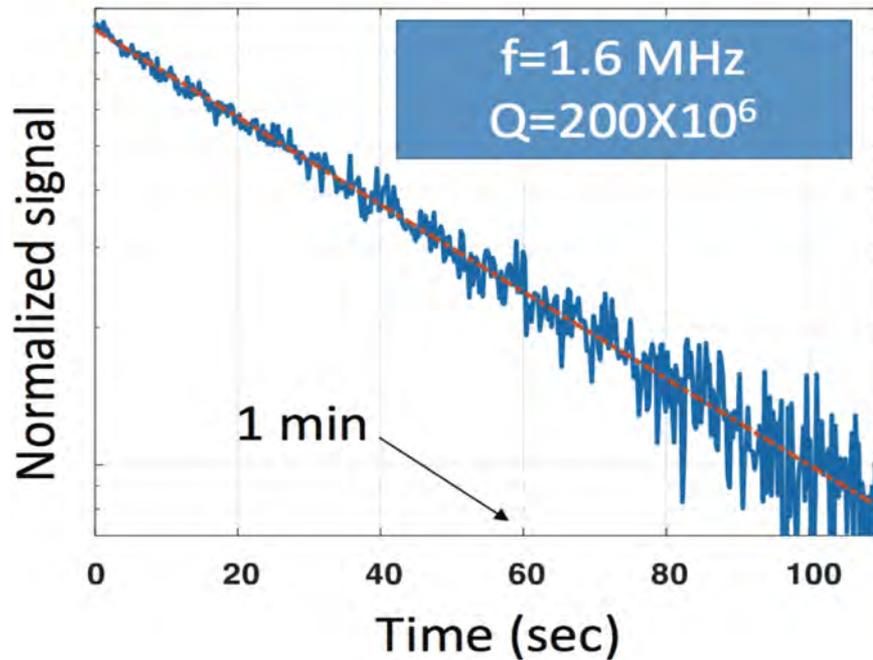


Mayer Alegre *et al.*, Optics Express (2011); Yu *et al.*, APL (2014)

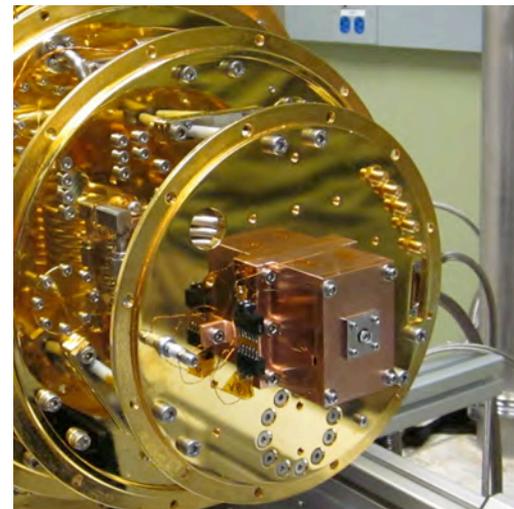
Y. Tsaturyan *et al.*, Nature Nanotech (2016); M. Yuan...Steele, APL (2015)

Extreme mechanical properties

Drive membrane and watch energy decay



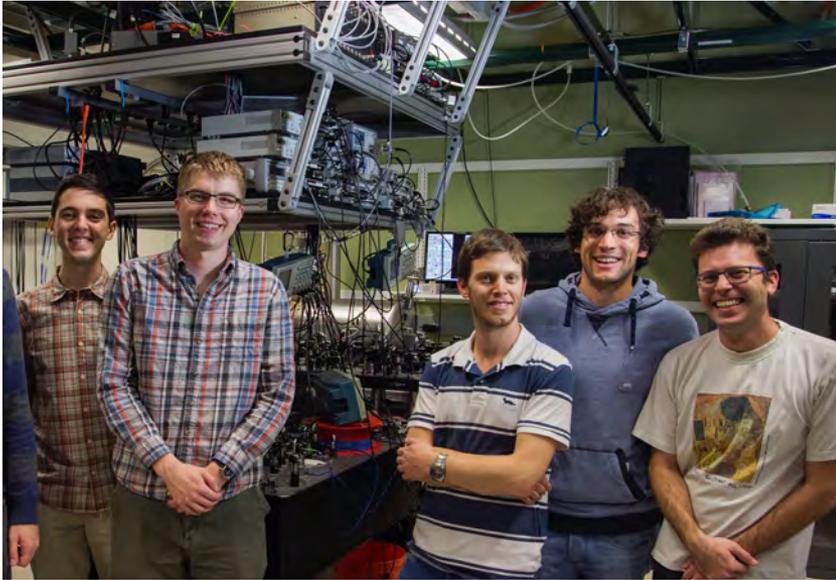
Corresponding heating rate = 10 quanta/ms



T=40 mK

The team

Regal group optomechanics team



Max Urmey

Bob Peterson Ran Fischer Nir Kampel
Gabriel Assumpcao (undergrad) Oliver Wipfli



DURIP

